

Are you being SERVed?
Data mining at the
coalface

Mark Lacy (“Science” hat!)

The SERVS survey

- * Spitzer warm mission survey of 18 deg^2 of sky to microJy sensitivity in the near-infrared (3.6-4.5 microns)
- * Laughably small data volume (few tens of GB) for SERVS alone
- * But modern astronomy a multiwavelength discipline, and we have multiwavelength complementary data from radio through X-ray wavelengths.
- * A real problem to make sense of objects with 10+ band detections over 10 decades of frequency.
- * Frequency over this range gives a whole new dimension in complexity.

SMALL DATA - 1 DEGREE, WITH $Z \sim 1$ GALAXY CLUSTERS FOUND USING VORONI TESSELATIONS (SIMPLE TOPOLOGICAL TECHNIQUE, GEACH ET AL 2011)

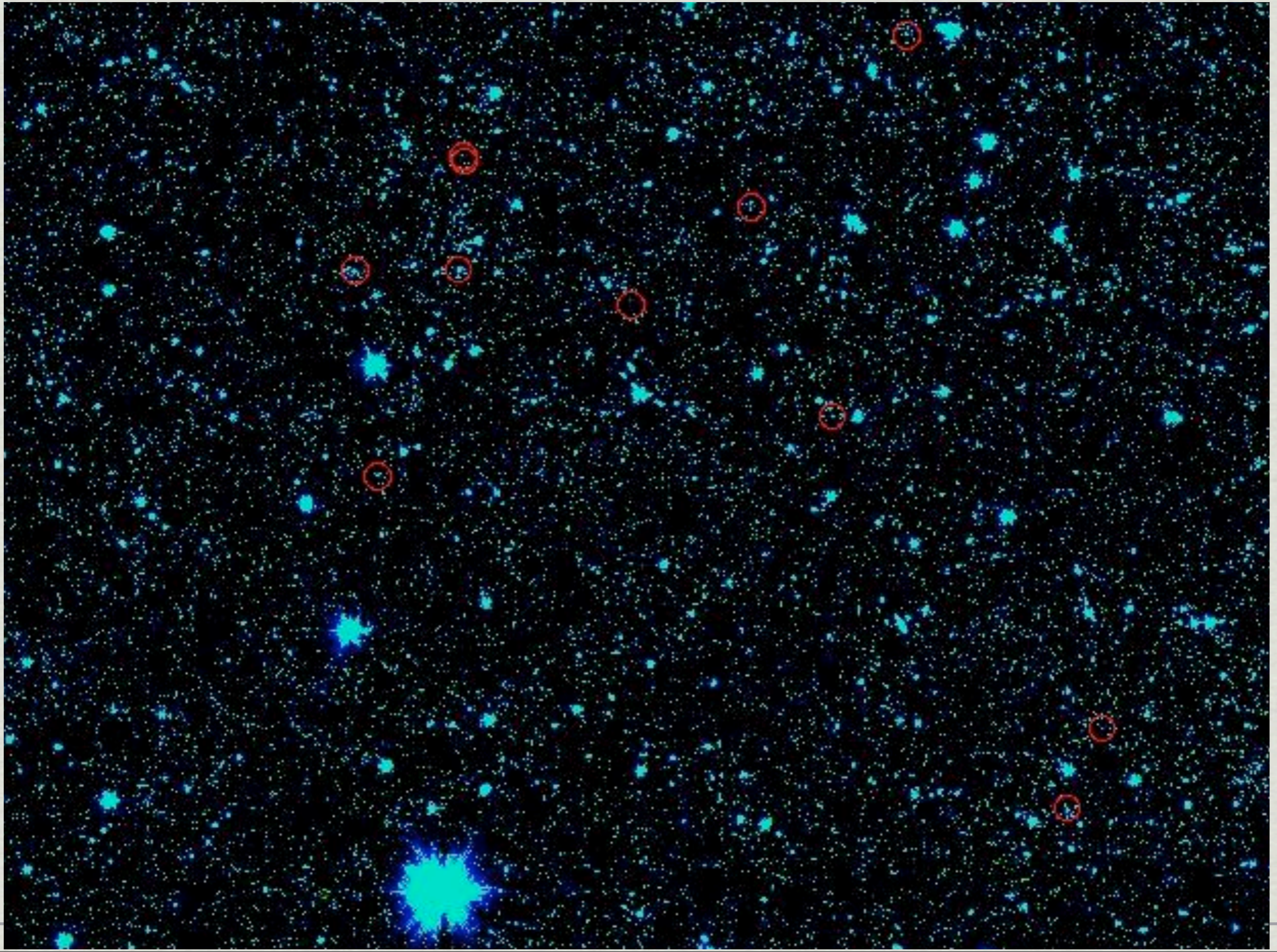
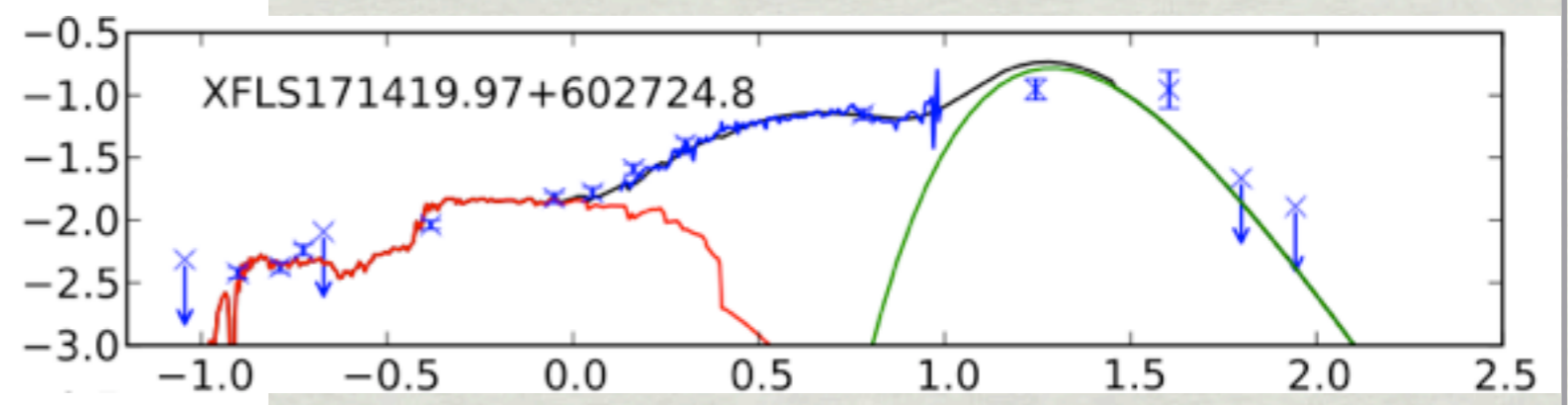
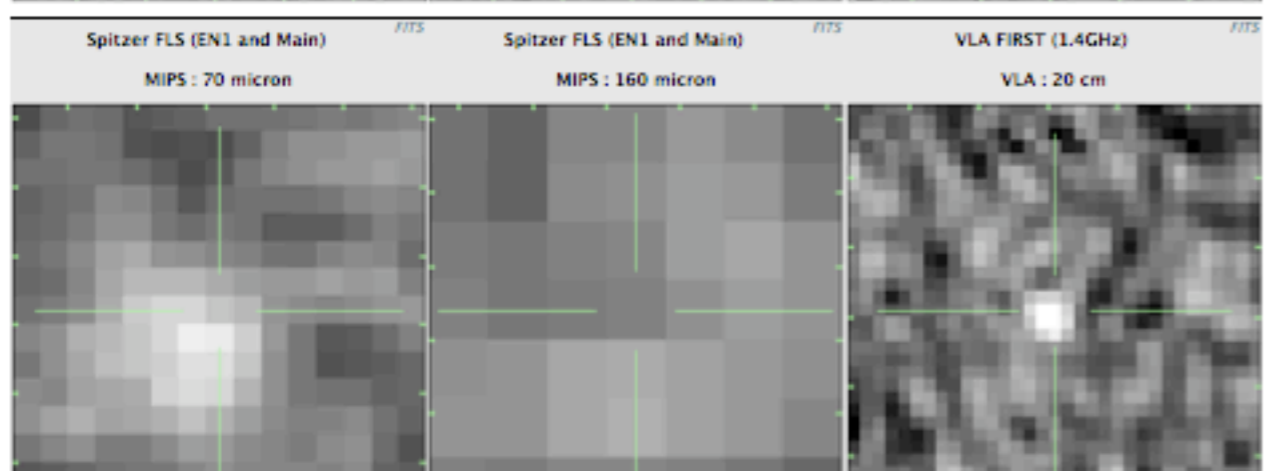
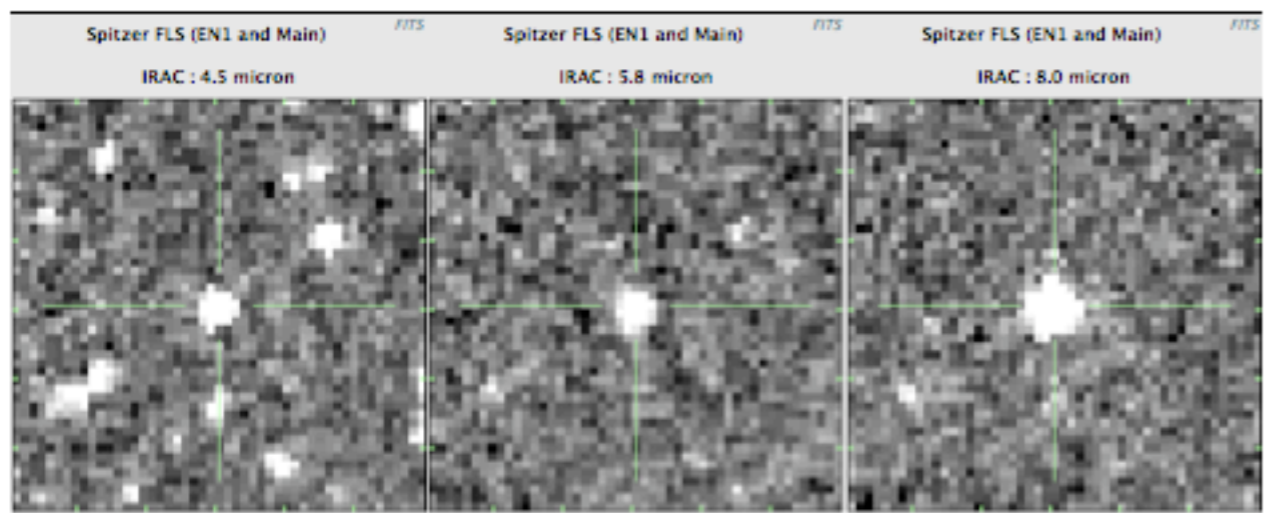
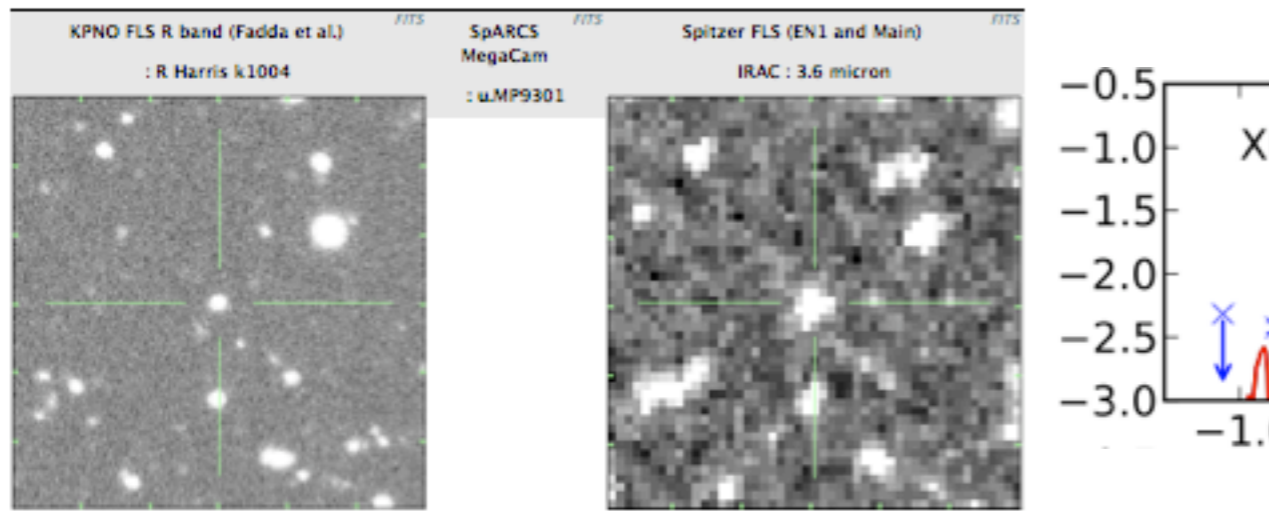
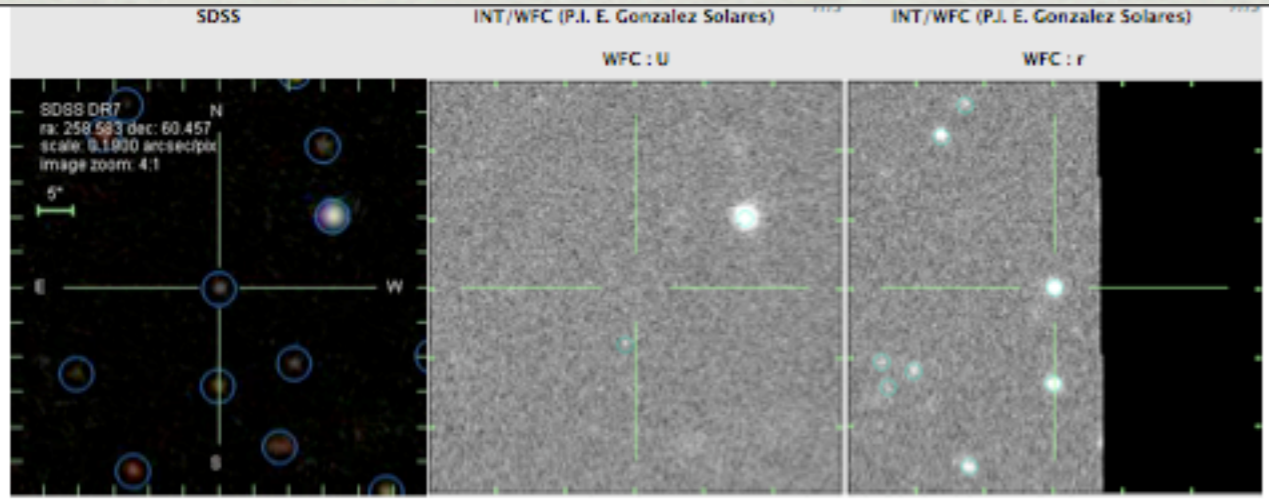
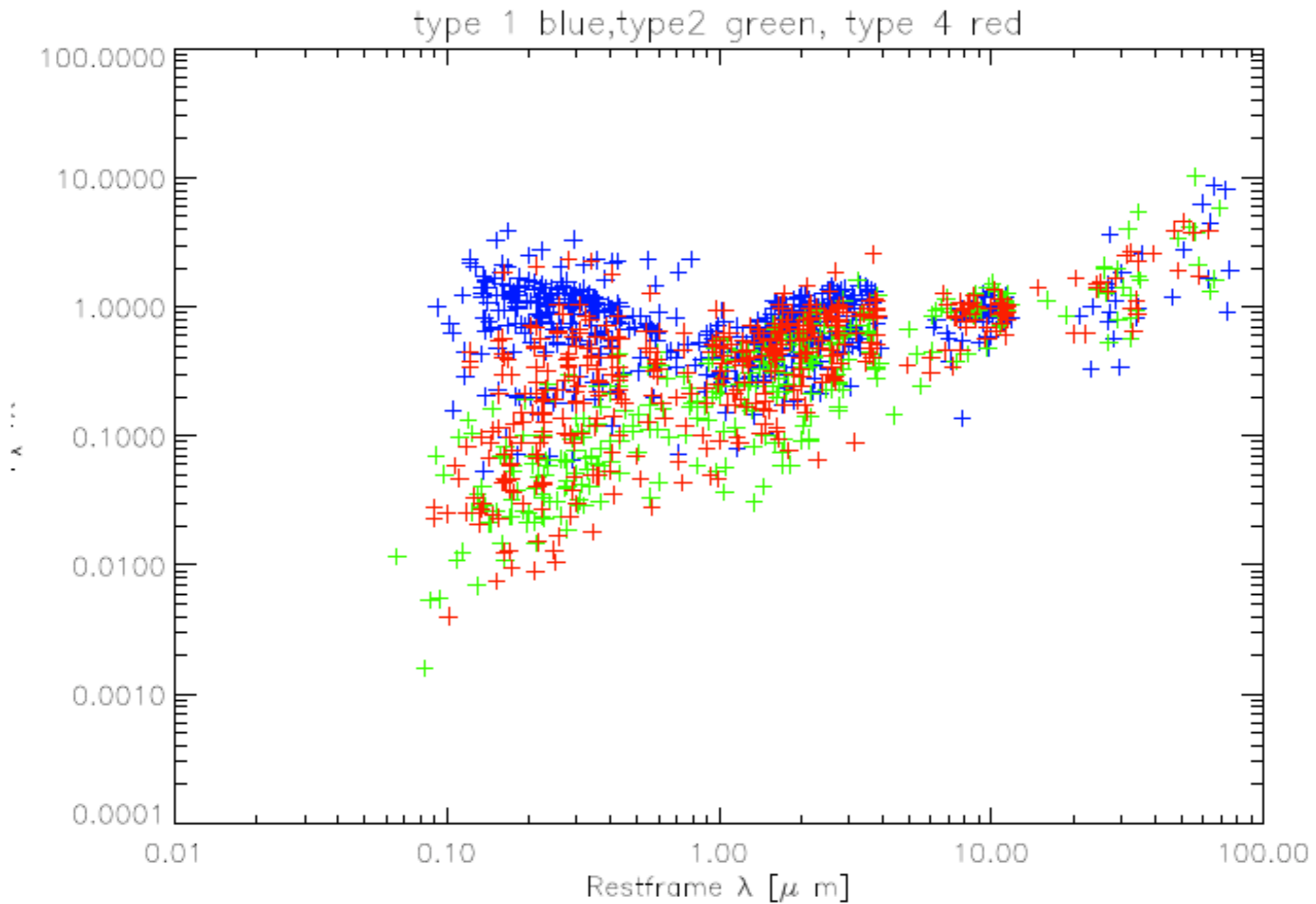


Image cutouts

- ✱ Eduardo Gonzales-Solares (Cambridge) has developed a simple cutout service to look at individual objects.
- ✱ A good illustration of max loading of brain...
- ✱ But very useful for spotting artifacts, determining whether or not an object is truly detected etc.

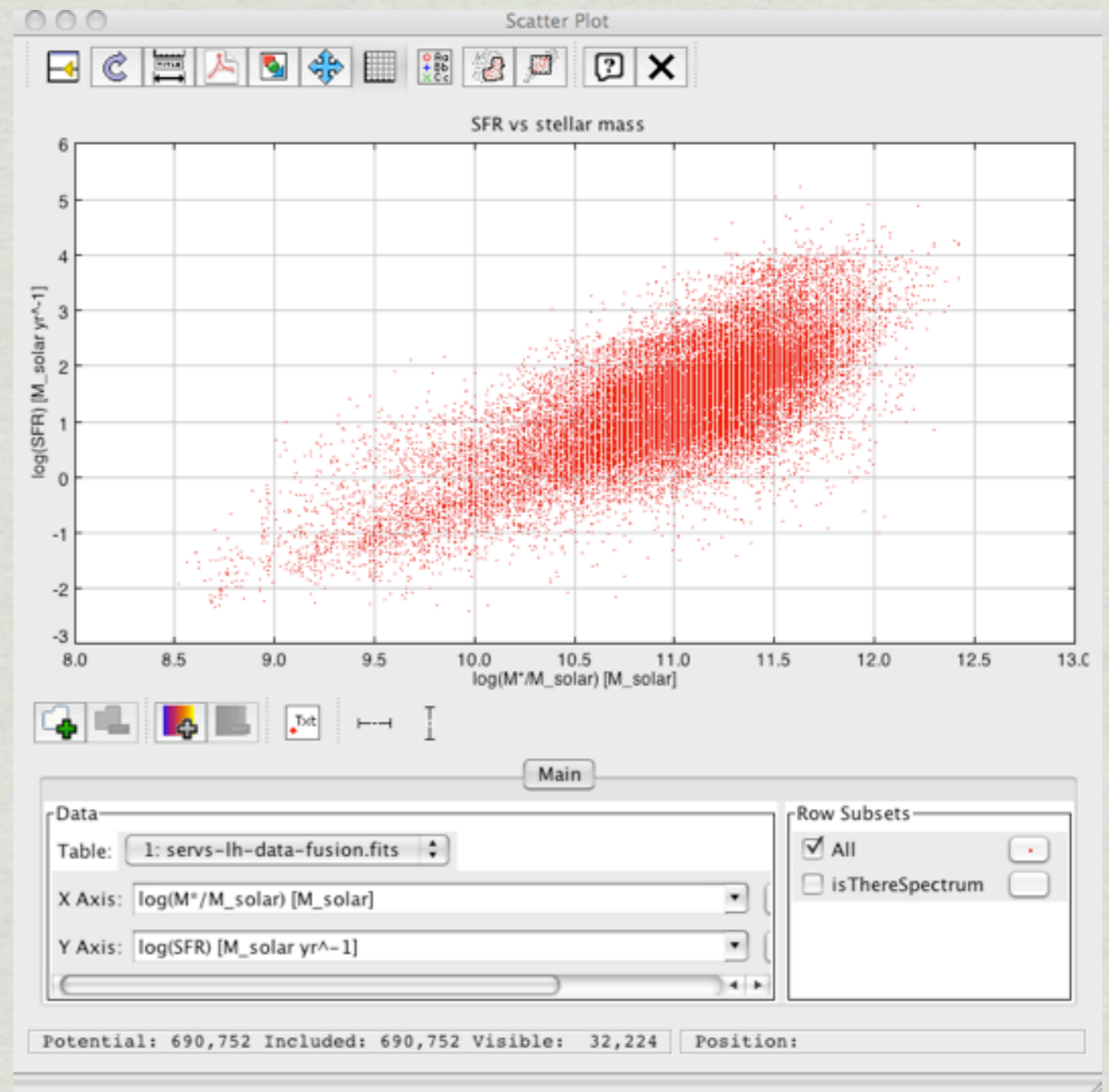


PARALLEL COORDINATES (PETRIC ET AL. IN PREPARATION)



Data fusion (Mattia Vaccari, Lucia Marchetti, Padova)

- * Merging catalogs from all these surveys (up to ~10 per field) gives ~1000 column, multi-GB merged catalog for a square degree.
- * Client applications such as Topcat can just deal with these now.
- * Extremely powerful datasets if they can be understood.



How do we go forward?

- * Catalogs already almost too large for desktops - may need to move to server-side apps fed by VO queries.
- * Association algorithms across 10 orders in frequency, widely varying resolutions (spatial and spectral), some surveys confusion limited, are difficult (e.g. GOODS Tfit).
- * High quality input catalogs are essential as you can't check every data point (or even 1% of them). Particularly important for rare object searches (e.g. distant quasars).
- * Software to provide quick, robust object characterizations (photo-zs, stellar masses, AGN/starburst breakdown).
- * Limits as important as detections!