

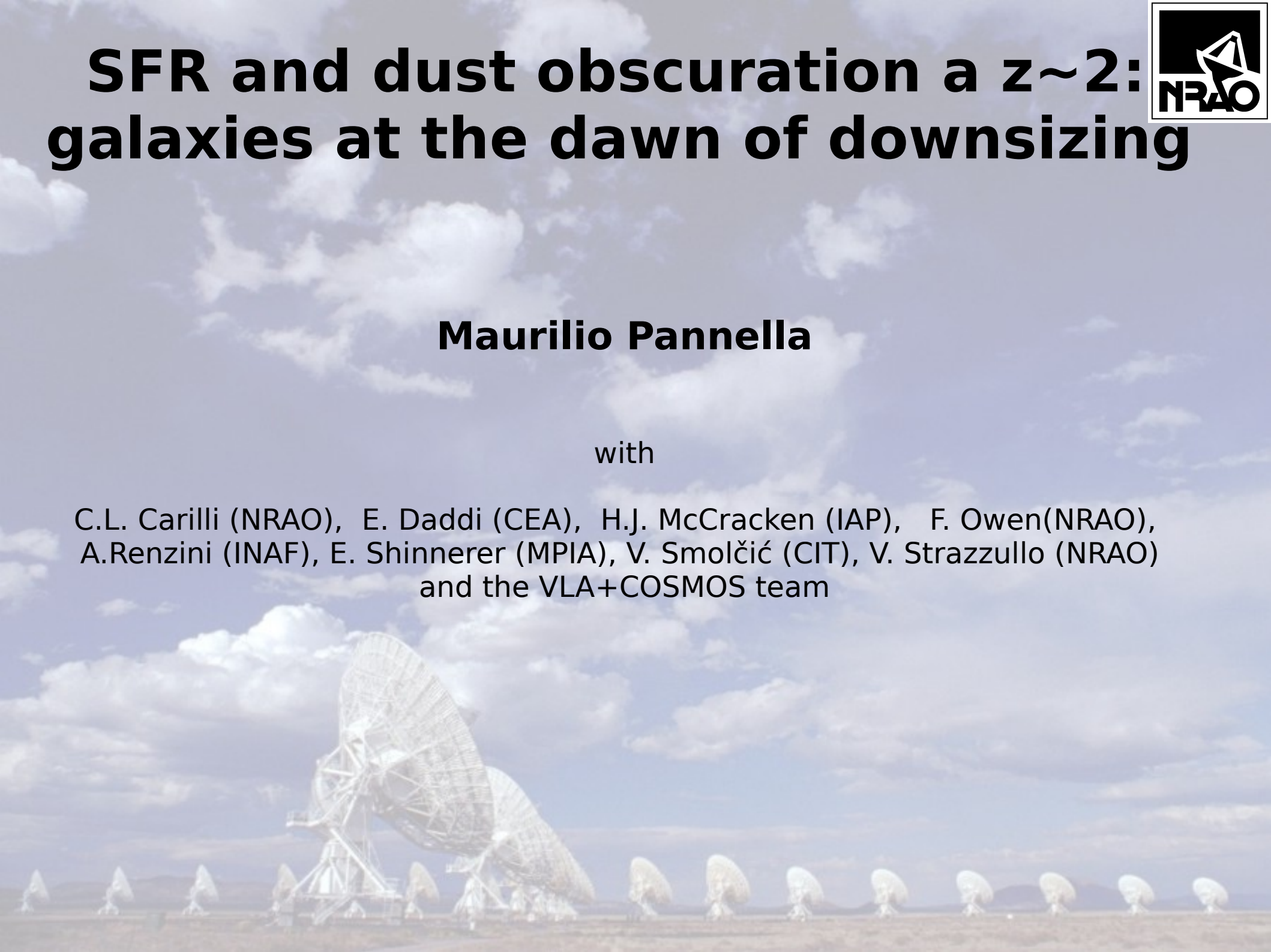
# **SFR and dust obscuration at $z \sim 2$ : galaxies at the dawn of downsizing**



**Maurilio Pannella**

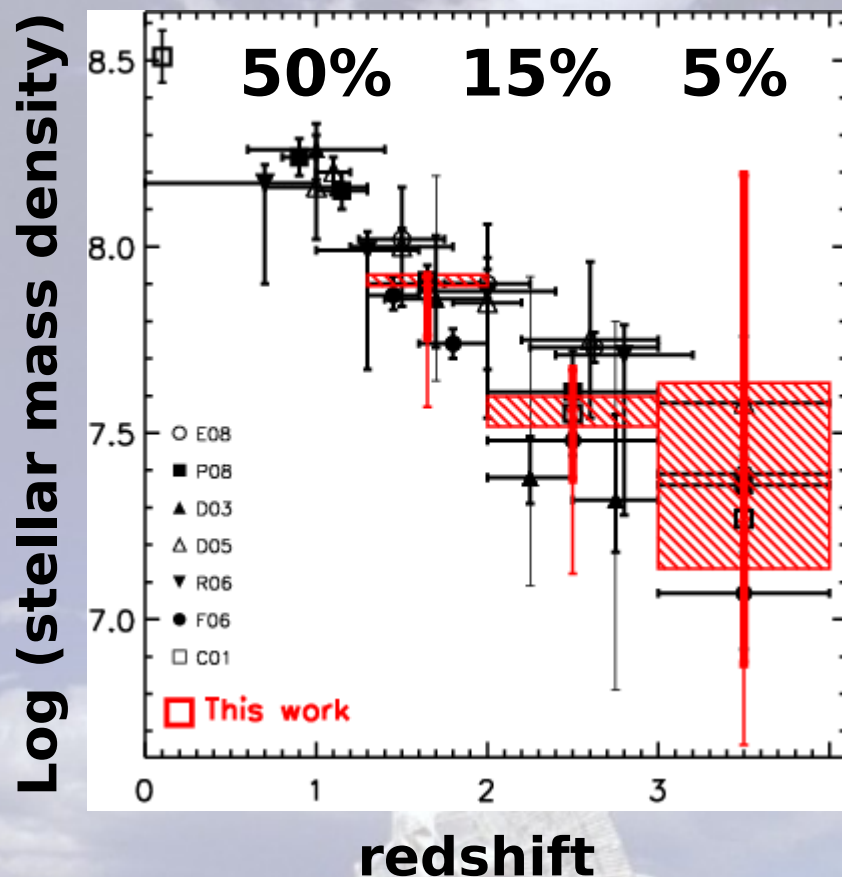
with

C.L. Carilli (NRAO), E. Daddi (CEA), H.J. McCracken (IAP), F. Owen (NRAO),  
A. Renzini (INAF), E. Shinnerer (MPIA), V. Smolčić (CIT), V. Strazzullo (NRAO)  
and the VLA+COSMOS team



# When and how galaxies formed

## The growth of stellar mass in the Universe



(Marchesini et al., 2009)

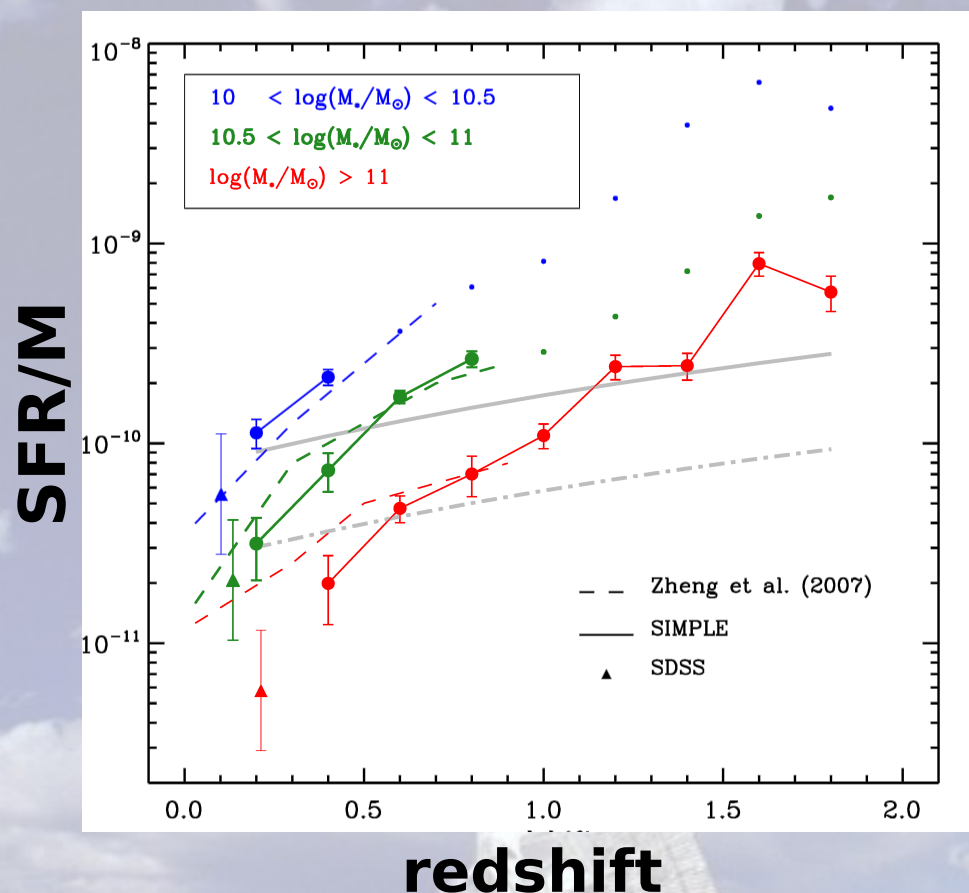
Broad consensus on the evolution of the galaxy stellar mass function up to high redshift

About 45% of the present day stellar mass has been produced in about 3.6 Gyrs at  $1 < z < 3$

The remaining 50% has formed in the last 7.5 Gyrs at  $1 < z < 0$

# When and how galaxies formed

## The downsizing of cosmic star formation



(Damen et al., 2009)

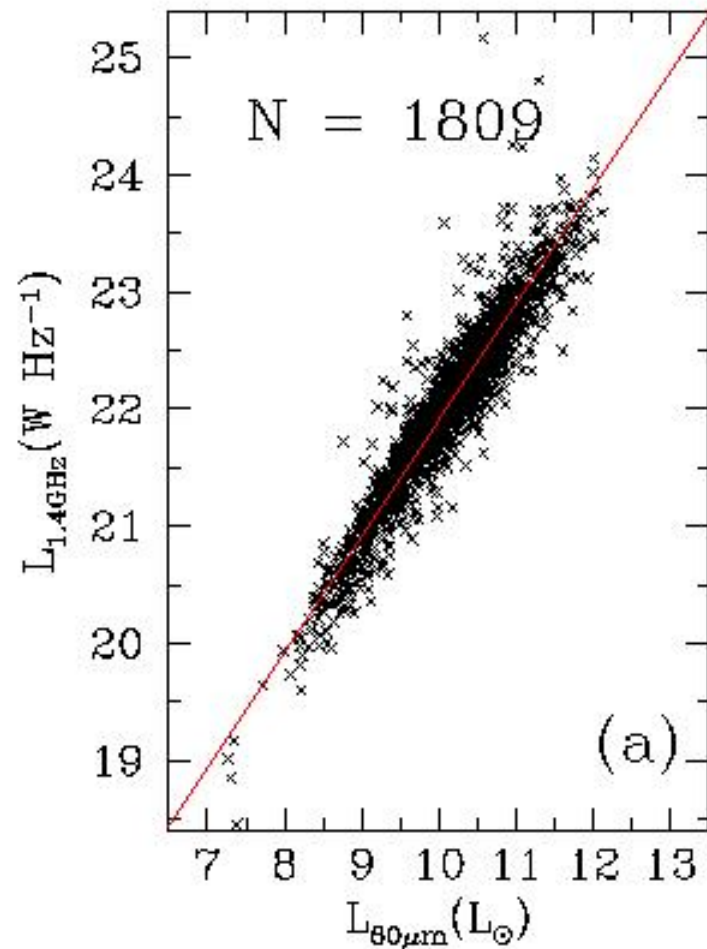
“The SSFR increases with  $z$  at a rate independent of mass”

“SSFRs of more massive galaxies are typically lower than those of less massive galaxies over the whole redshift range”

The downsizing pattern seems to be at work up to high redshift



# Radio emission and SFR



(Yun et al., 2001)

SFR-FIR correlation (Kennicutt 1998)

**x**

Radio-FIR correlation (Yun et al, 2001; Bell 2003)

**x**

Radio interferometry has  $\sim 1$  arcsec resolution

**=**

The ideal dust-unbiased SFR indicator

# The VLA-COSMOS wide survey



VLA Large Program

(P.I. Eva Schinnerer)

Full COSMOS field at 1.4 GHz

1.5" resolution

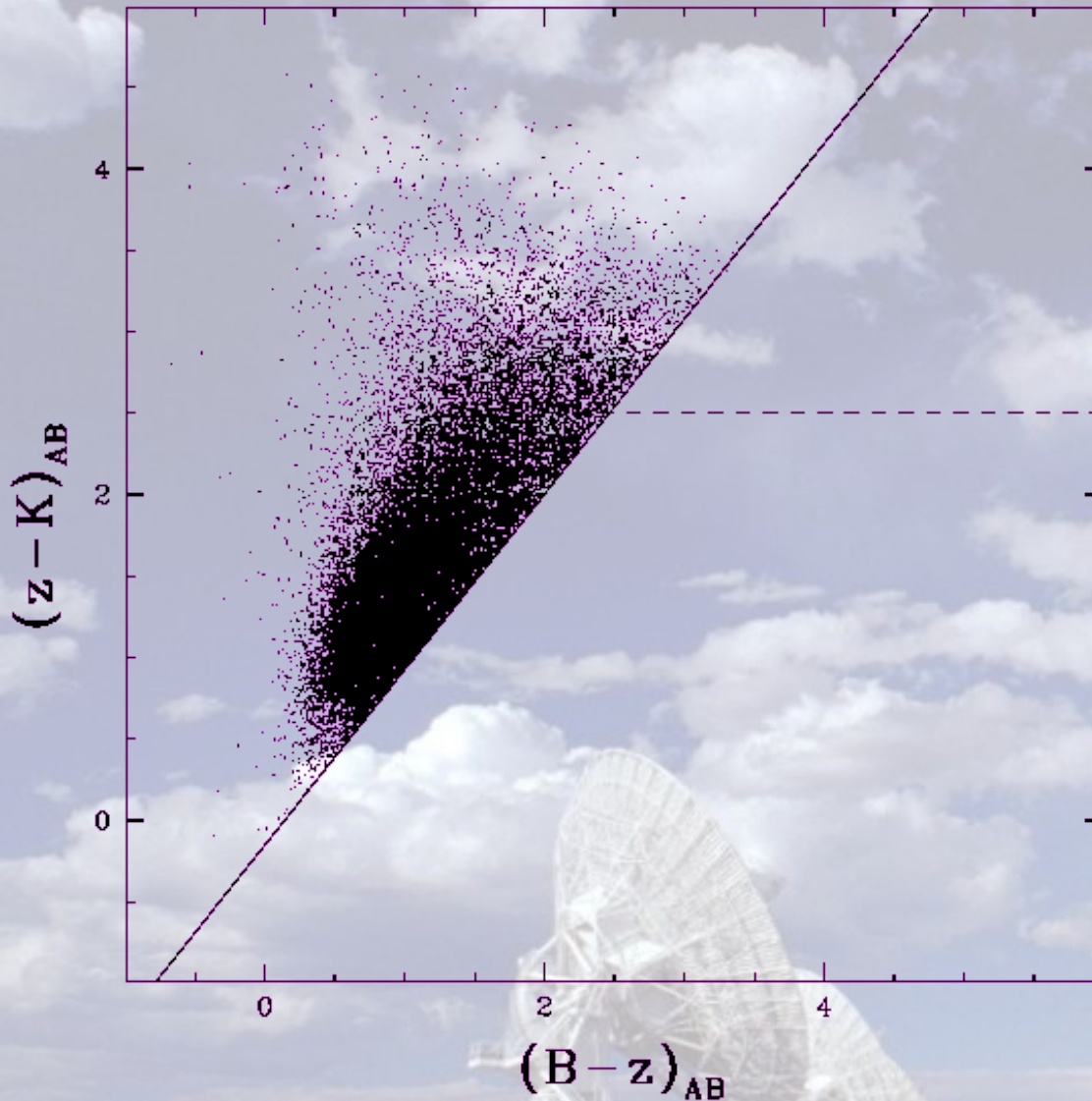
rms  $\sim 10 \mu\text{Jy}$



(Schinnerer et al., 2007)



# The BzK COSMOS project



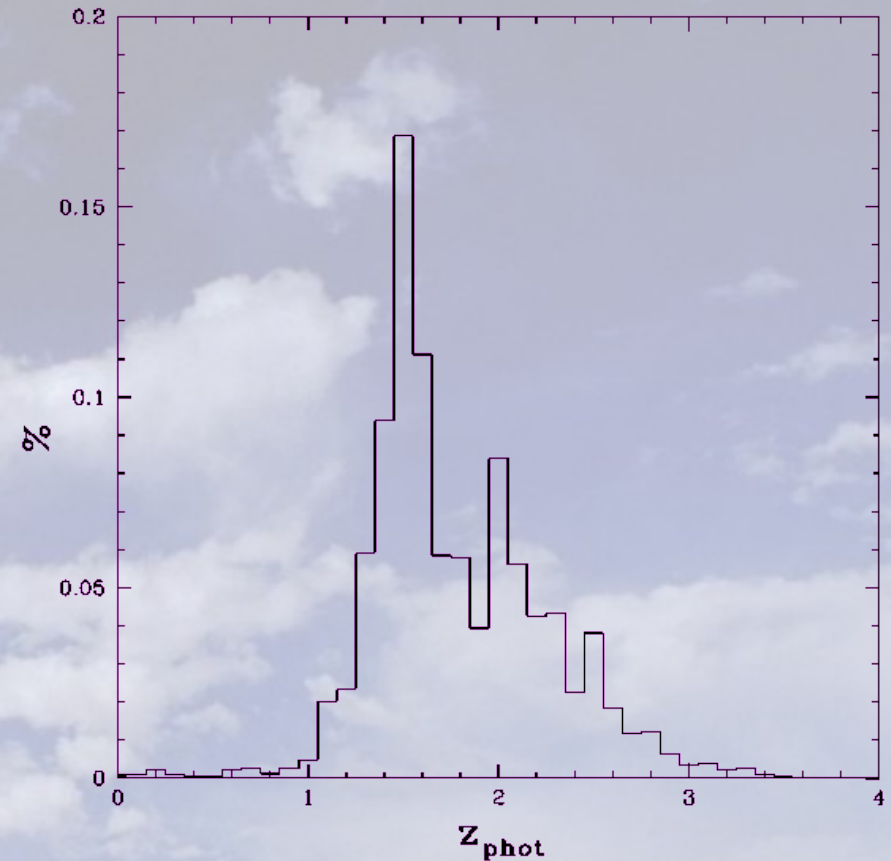
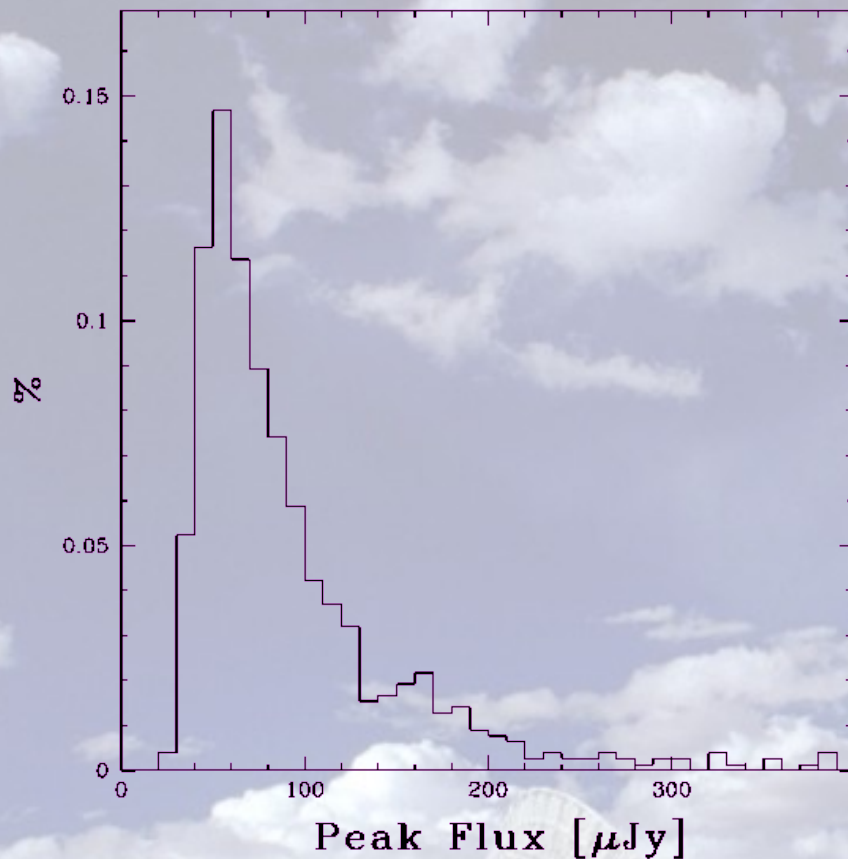
30125 sBzK galaxies

WIRCAM/CFHT K  
(P.I. H. J. McCracken)

SuprimeCam/Subaru Bz  
(COSMOS Legacy dataset)

$K_{AB} \sim 23$  mag

# The BzK COSMOS project



Extremely effective selection of galaxies at  $1.3 < z < 2.5$

“Only” 616 objects ( $\sim 2\%$ ) are 1.4 GHz detected

# The stacking analysis

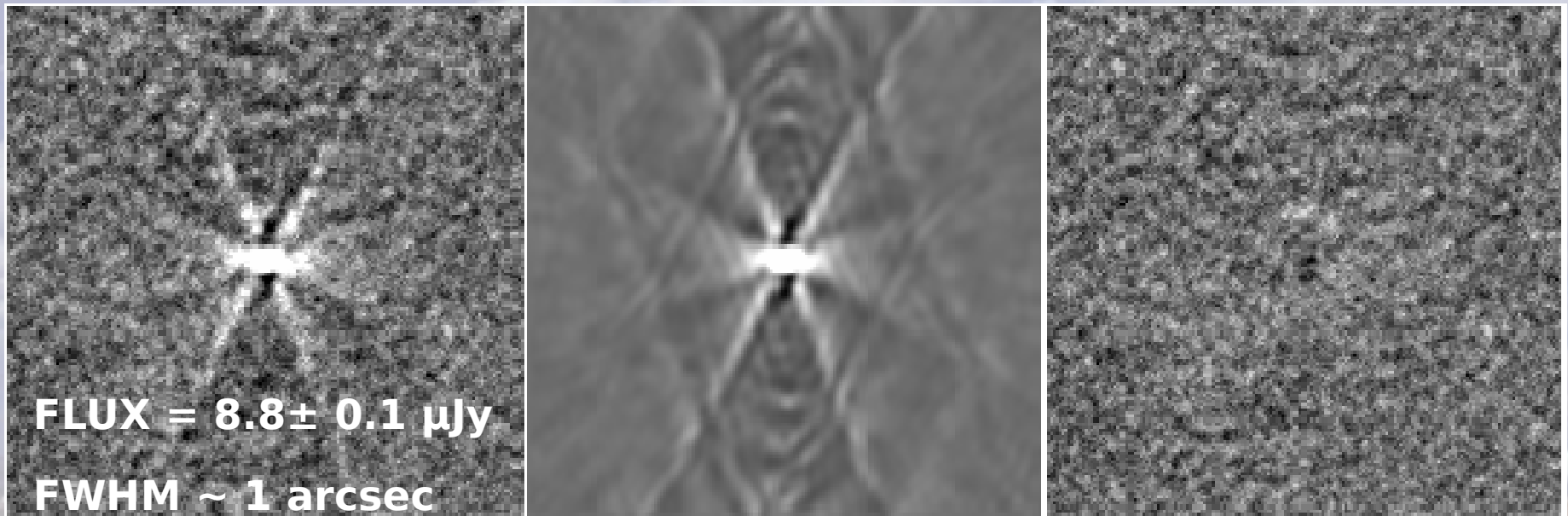
## 1.4 GHz median stacking:

- more robust than mean against detections
- rms goes down by  $\sim \sqrt{N}$  *i.e.* 0.1/0.3  $\mu\text{Jy}$
- “normal” star forming galaxy at high  $z$
- next generation arrays science case (SKA)

**Stacked data**

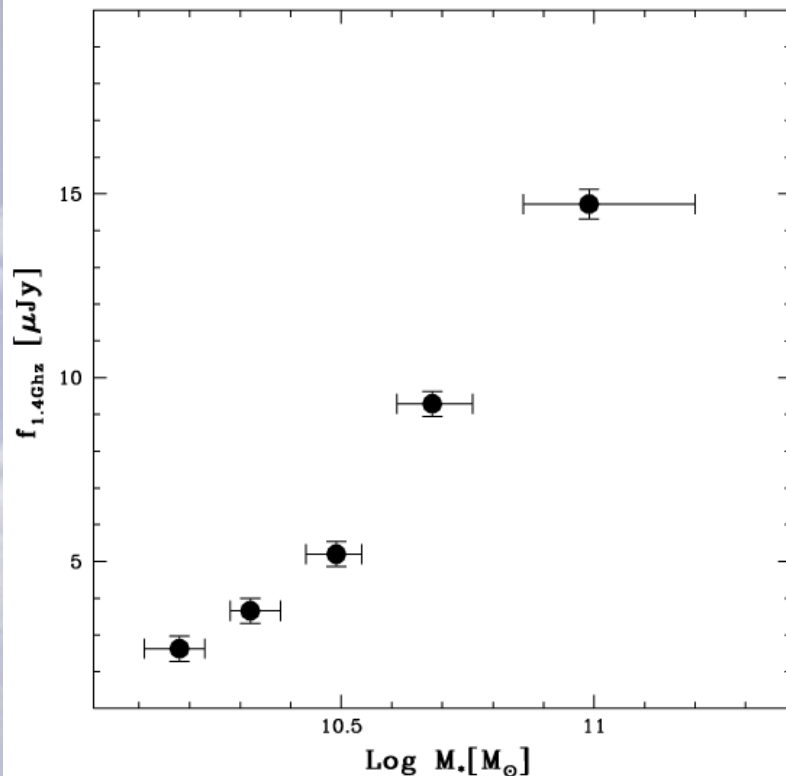
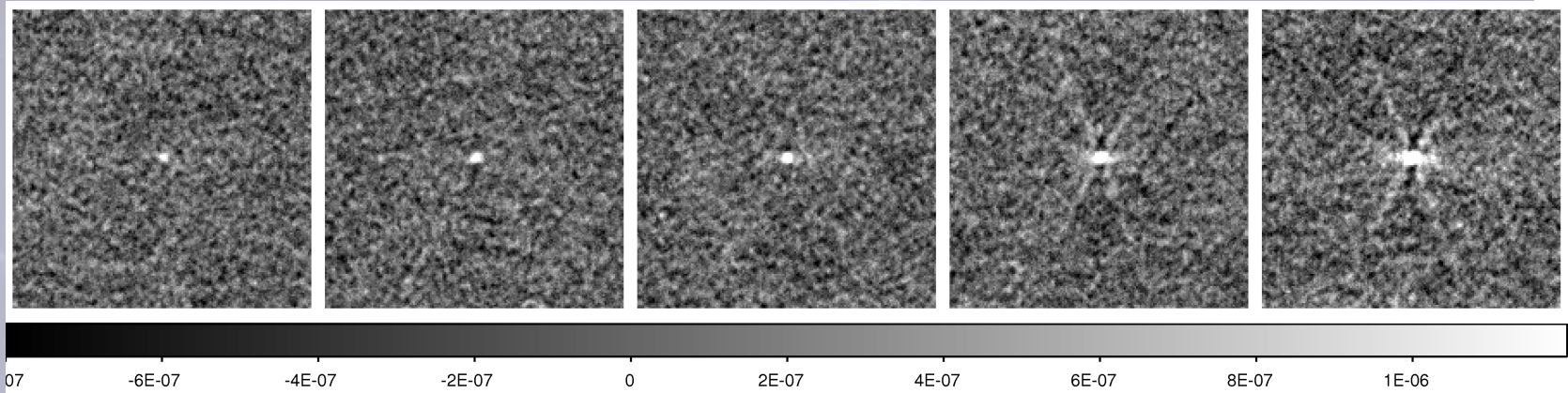
**Model**

**Residuals**



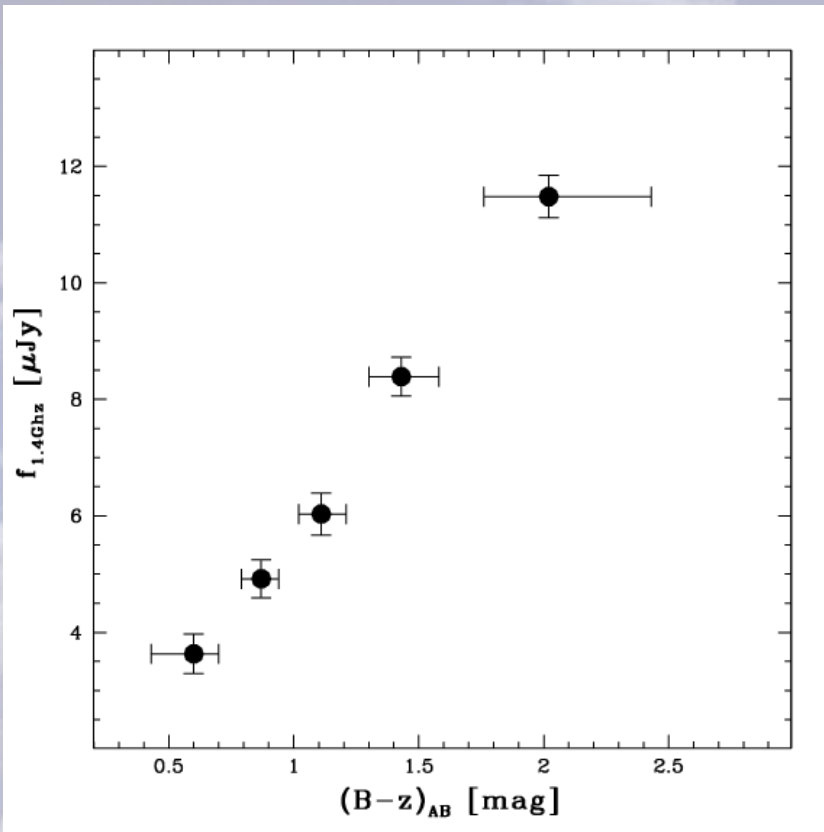
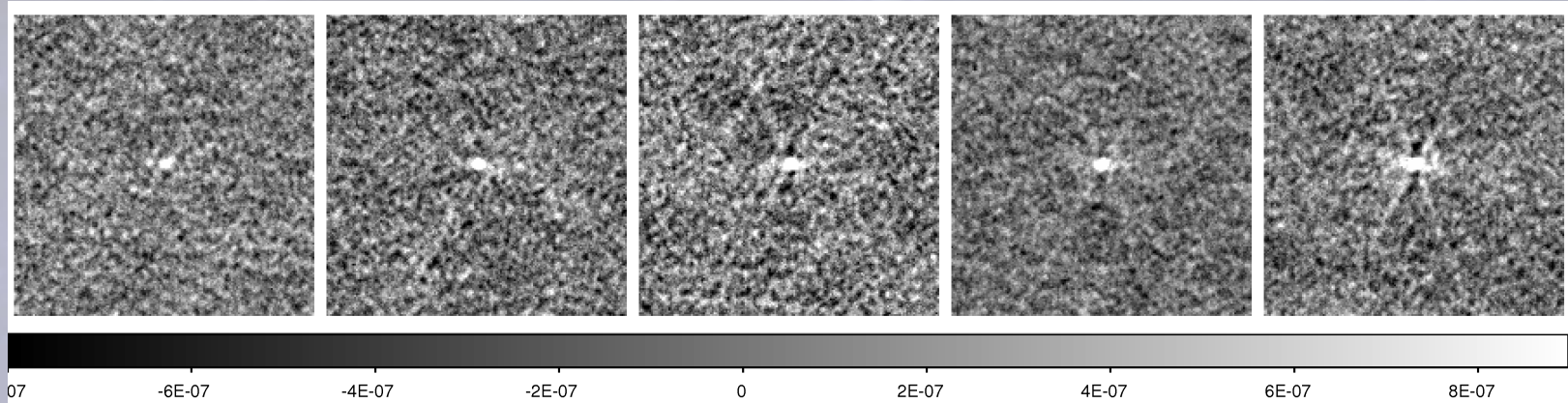


# Radio stacks vs. Stellar Mass



- Tight correlation between galaxy stellar mass and star formation
- Similarly to the local Universe: the higher the stellar mass, the more the star formation

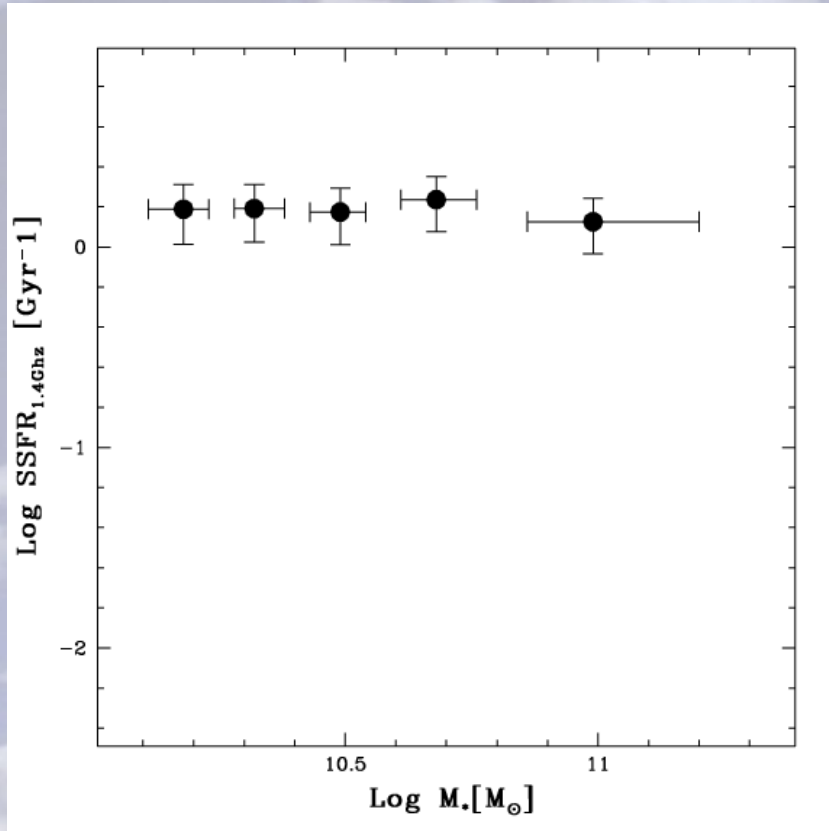
# Radio stacks vs. (B-z) color



- Tight correlation between (B-z) color and star formation activity
- The observed (B-z) color of  $z \sim 2$  star forming galaxies is a measure of the UV slope, *i.e.* the dust content

**Galaxies with higher SFRs  
are more dust extinguished**

# Radio SSFRs



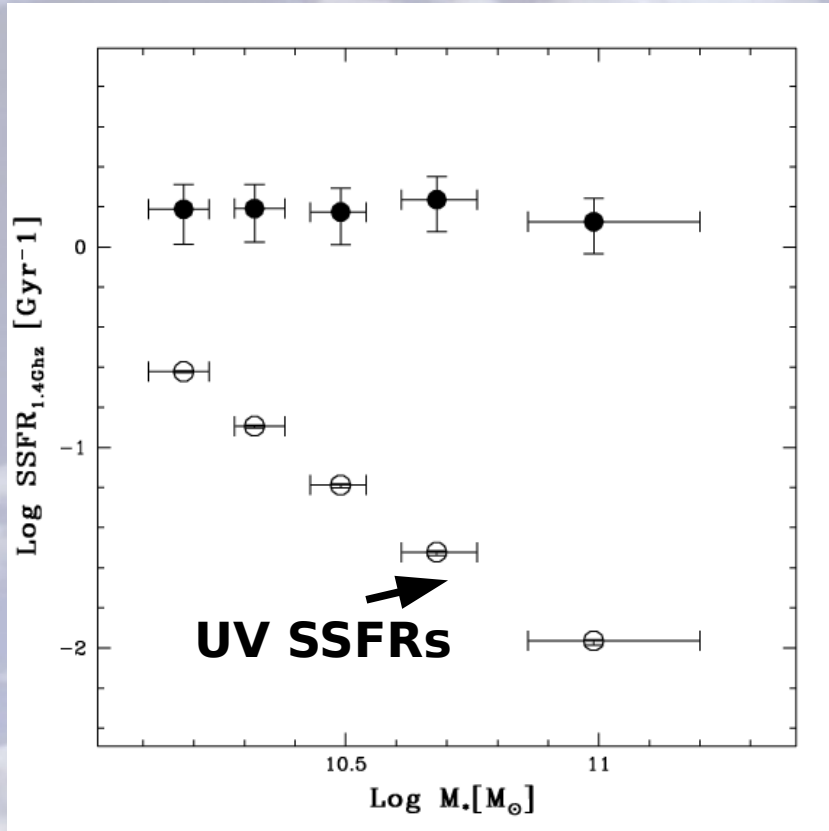
SSFR vs stellar mass is  $\sim$ flat at  $z = 1.7$

Galaxies are all in their active epoch

The mass growth is the same at all  $M_*$

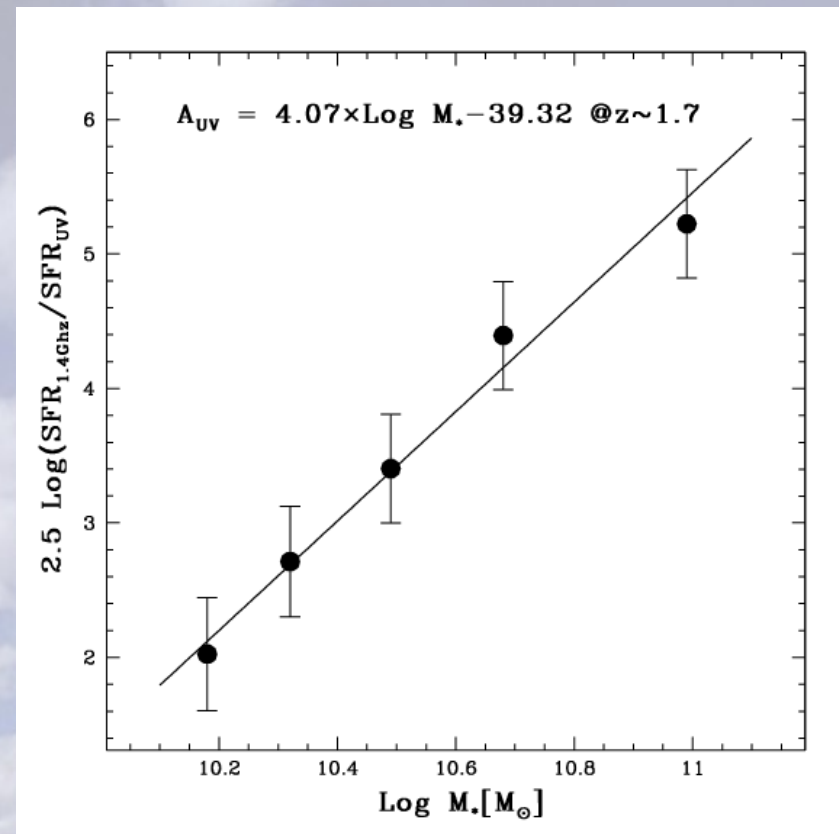
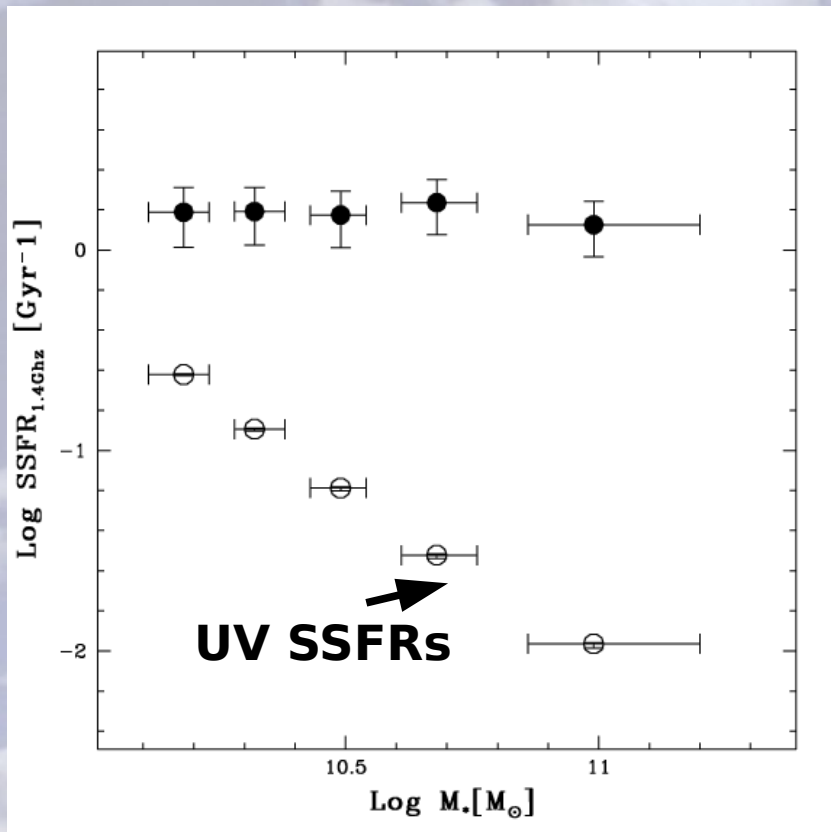


# Radio vs UV SSFRs



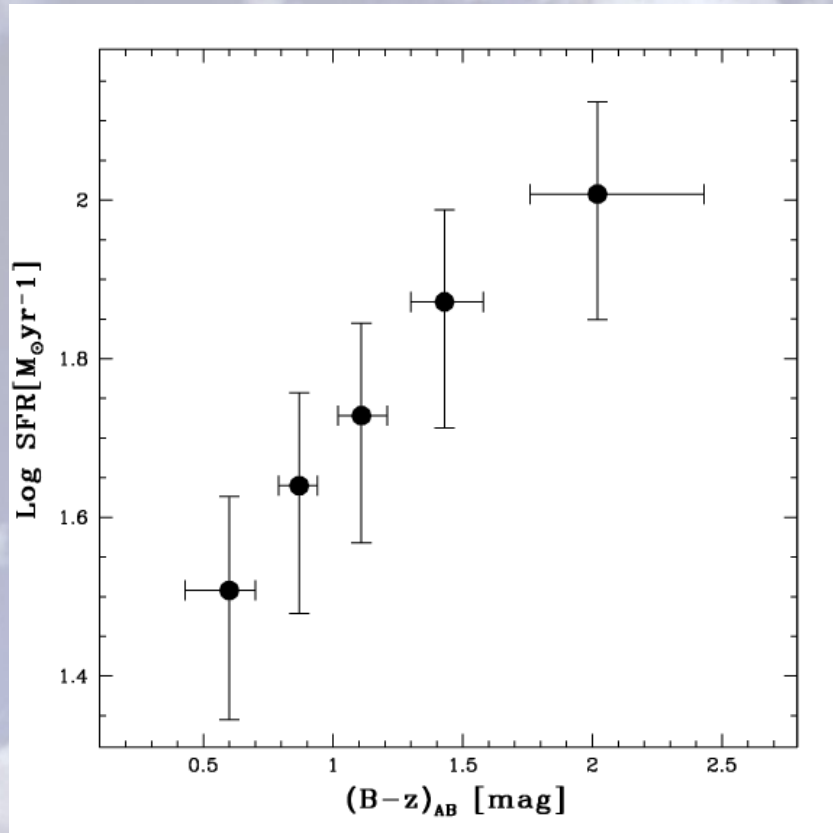
**Massive starburst galaxies are as red as ETGs**

# Radio vs UV SSFRs



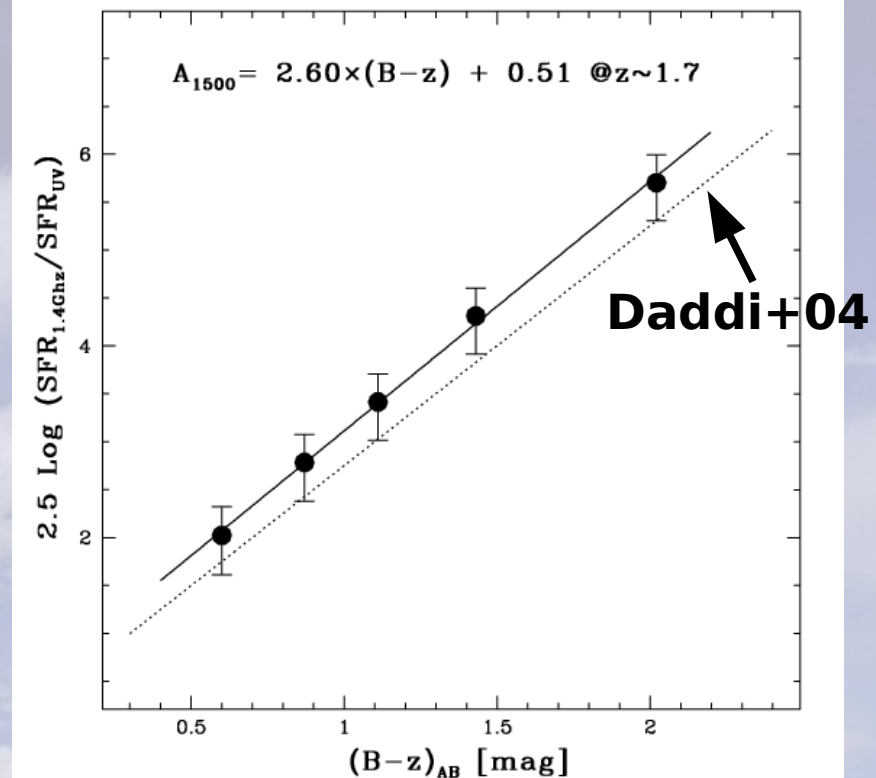
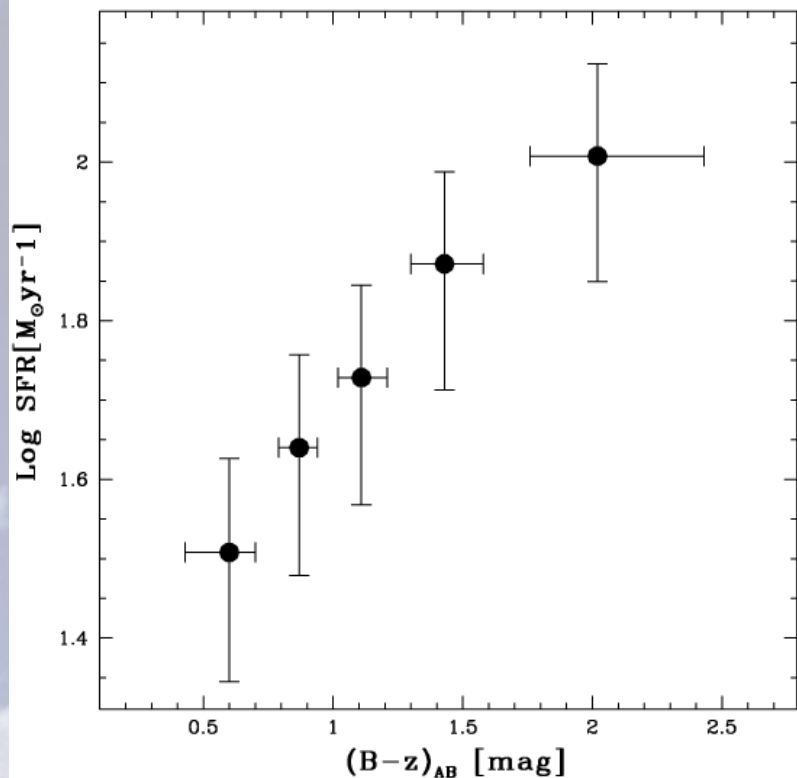
**Massive starburst galaxies are as red as ETGs**  
 **$A_{1500}$  is linearly proportional to  $\text{Log } M_*$**

# Radio SFRs vs (B-z)



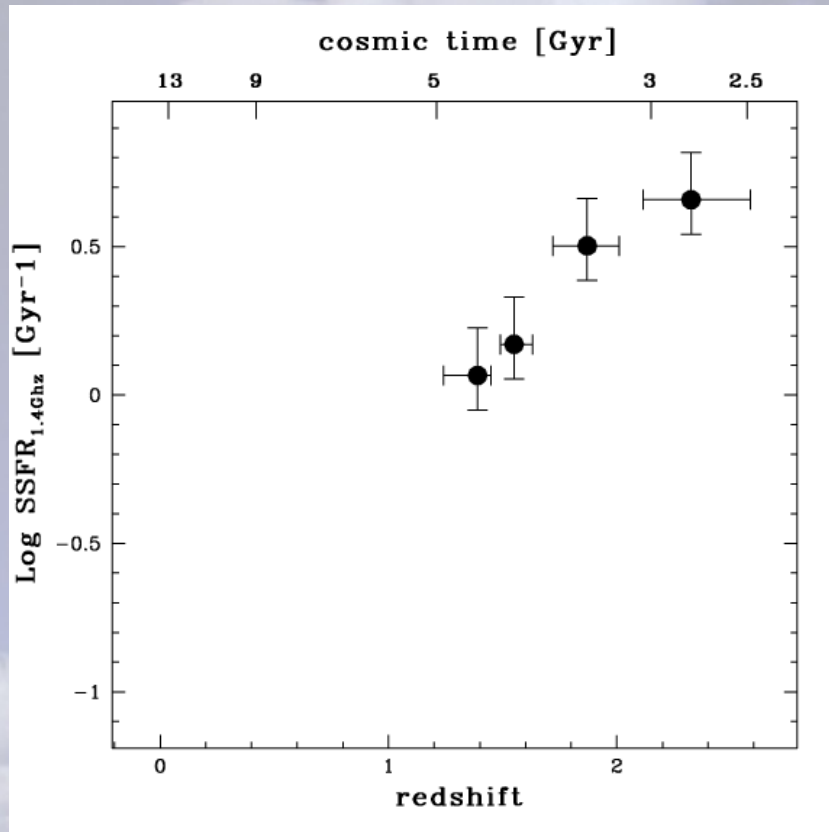


# Radio SFRs vs (B-z)



**$A_{1500}$  is a linear combination of  $\text{Log } M_*$  and  $(B-z)$**

# The SSFR redshift evolution



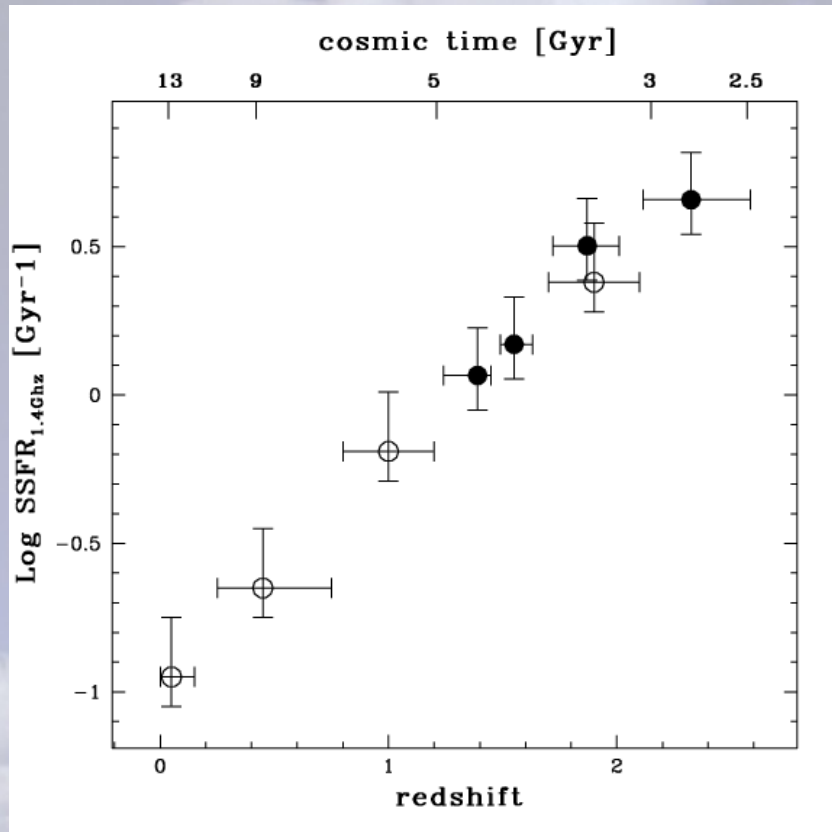
**SSFR(M) is flat in  $1.4 < z < 2.3$**

**SSFR(z) decreases by a factor 4 in the redshift range  $1.4 < z < 2.3$**

# The mass growth of galaxies



## The secular decline of SSFR with time



**SSFR(M) is flat in  $1.4 < z < 2.3$**

**SSFR(z) decreases by a factor 4 in the redshift range  $1.4 < z < 2.3$**

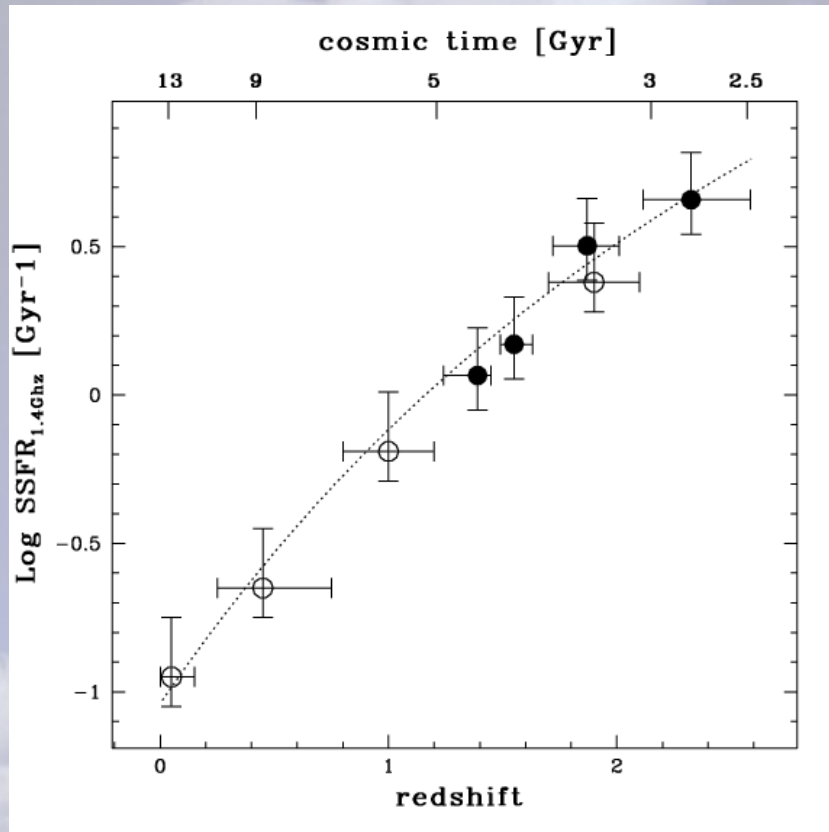
**The decrease continues all the way down to the local Universe**



# The mass growth of galaxies



$$\langle \text{SFR}(M,t) \rangle \approx 270 (M/10^{11}) (t/3.4)^{-2.5} = dM/dt$$



**Galaxies cannot keep growing at the empirical average SFR!**

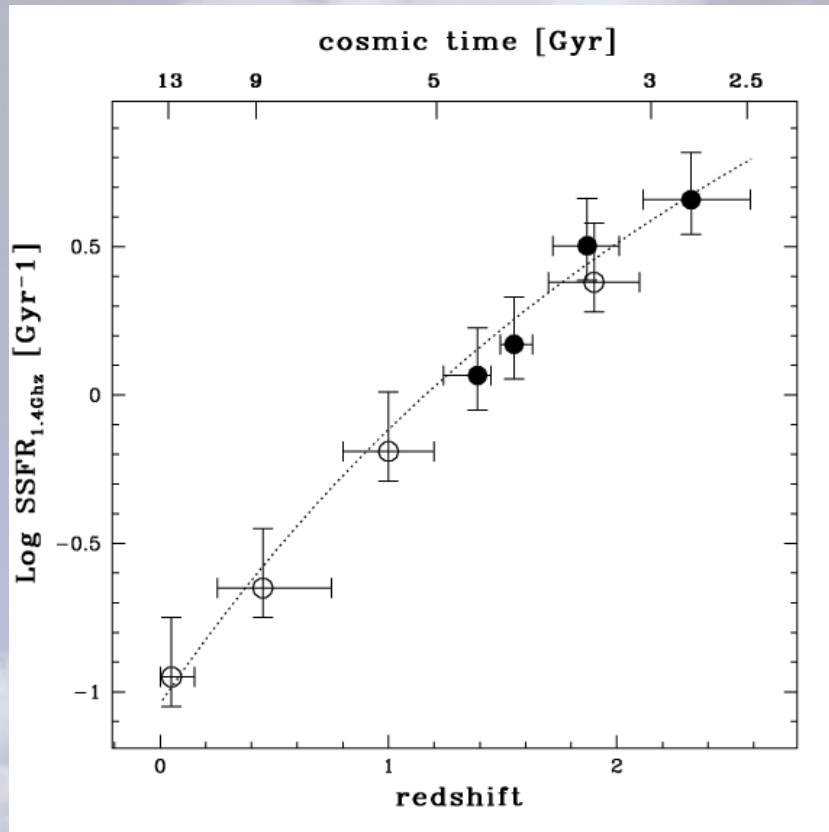
**If so, the mass of individual galaxies should increase by a factor ~6 between  $z=2$  and  $z=1.6$ , a factor ~20 by  $z=1.3$ , and ~250 by  $z=0$ .**

**Clearly, galaxies don't grow that much!**

# The mass growth of galaxies



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**If so, the mass of individual galaxies should increase by a factor  $\sim 6$  between  $z=2$  and  $z=1.6$ , a factor  $\sim 20$  by  $z=1.3$ , and  $\sim 250$  by  $z=0$ .**

**Clearly, galaxies don't grow that much!**

**They turn passive  
They do so in a "downsized" fashion**

# Conclusions



**A universal dust attenuation correction does not apply**

**Dust attenuation is a function of galaxy stellar mass, with more massive galaxies being more heavily extinguished**

**Galaxies have, at all masses, the same evolutionary timescales and a nearly exponential growth with time**

**The mass overgrowth is not happening because galaxies turn passive in a downsized fashion**

**This might be regarded as the “dawn of downsizing” for the star forming galaxy population at  $z \sim 2$**

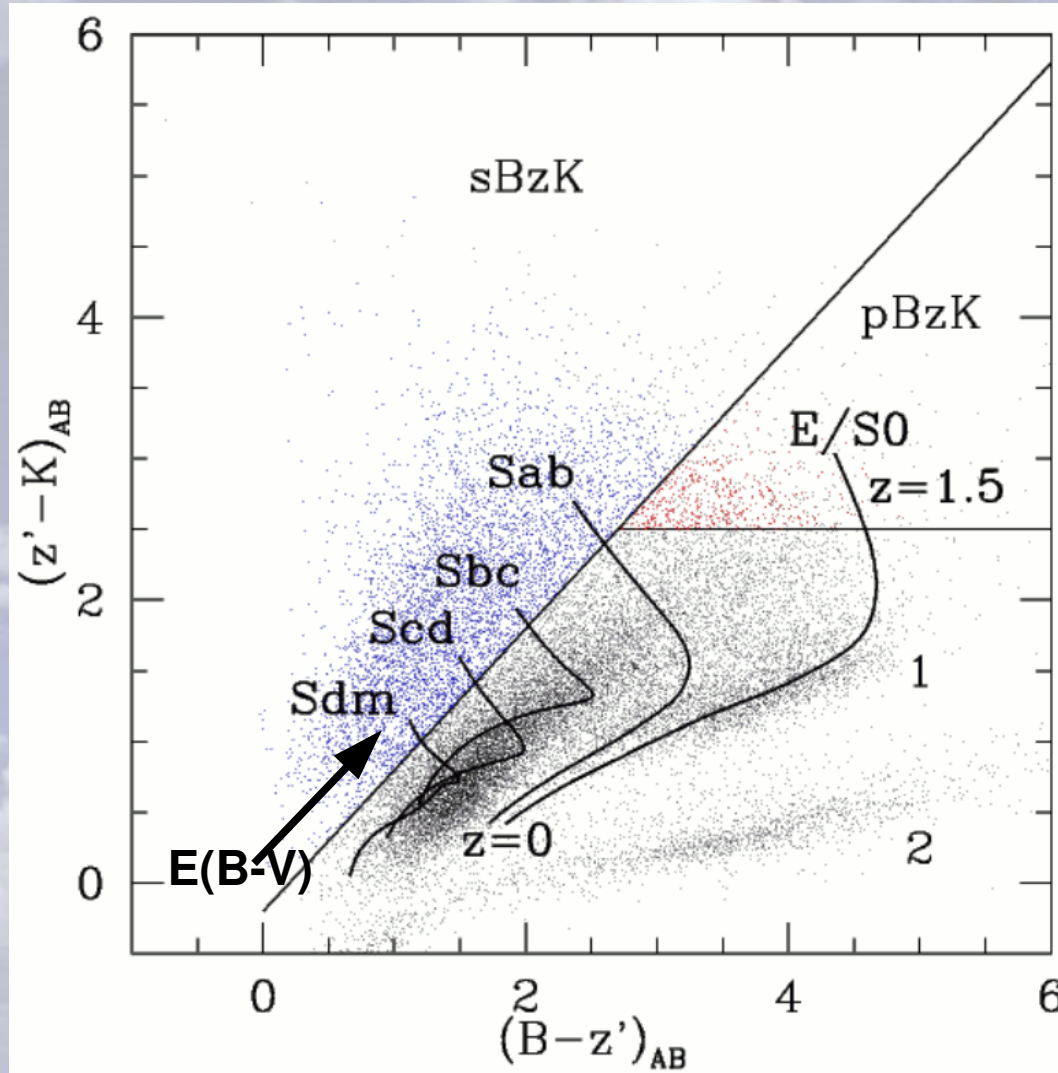
*“A different approach to galaxy evolution”* **from Alvio Renzini**



# The BzK COSMOS project



Chasing galaxies at  $z \sim 2$  : the BzK selection technique



(Lane et al. 2007)

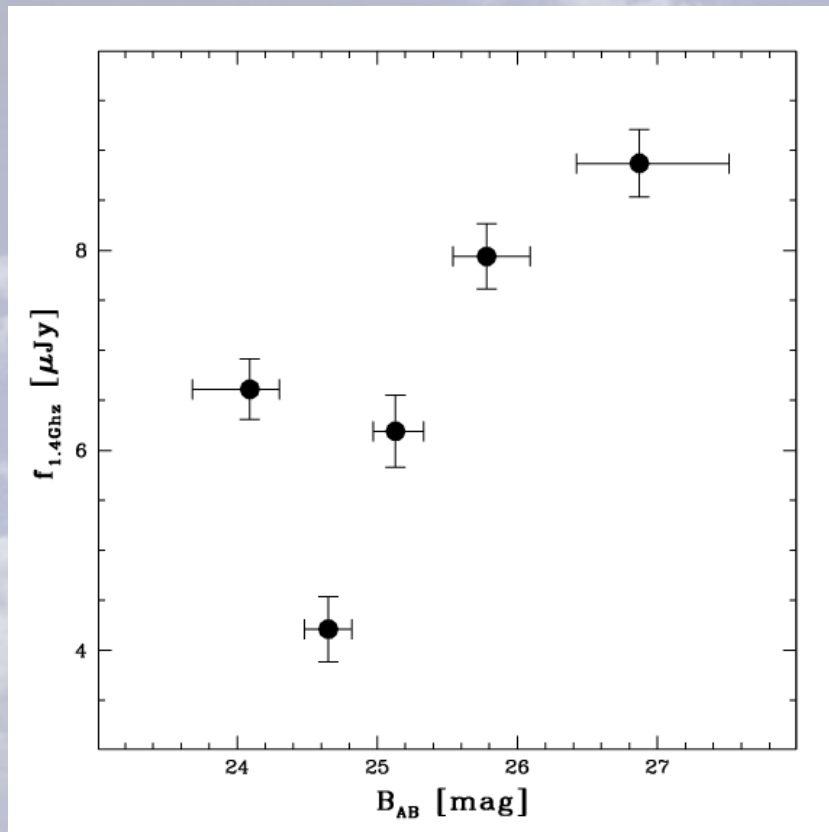
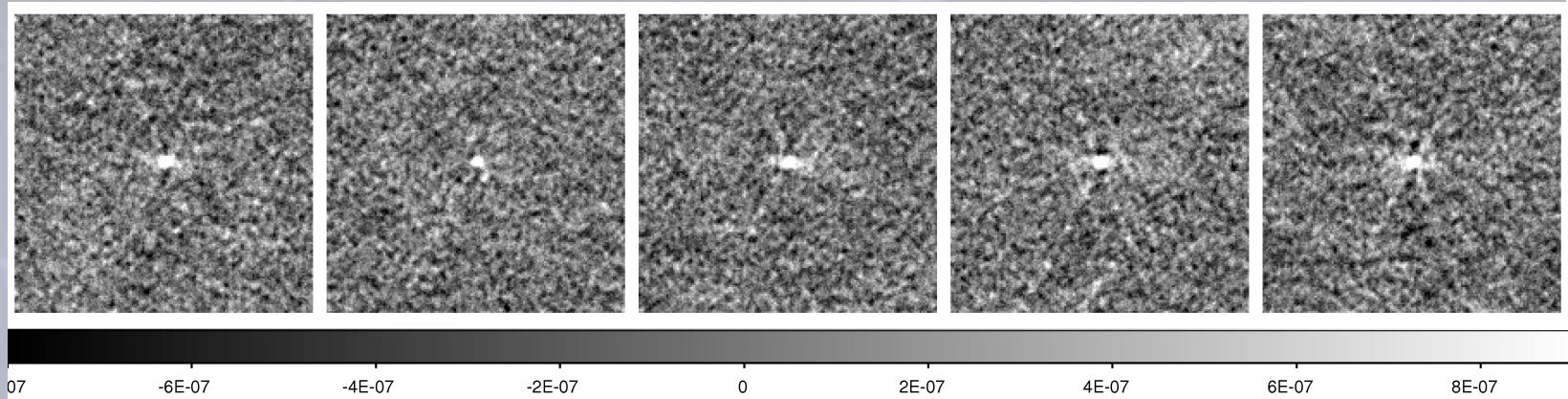
very simple

reddening independent

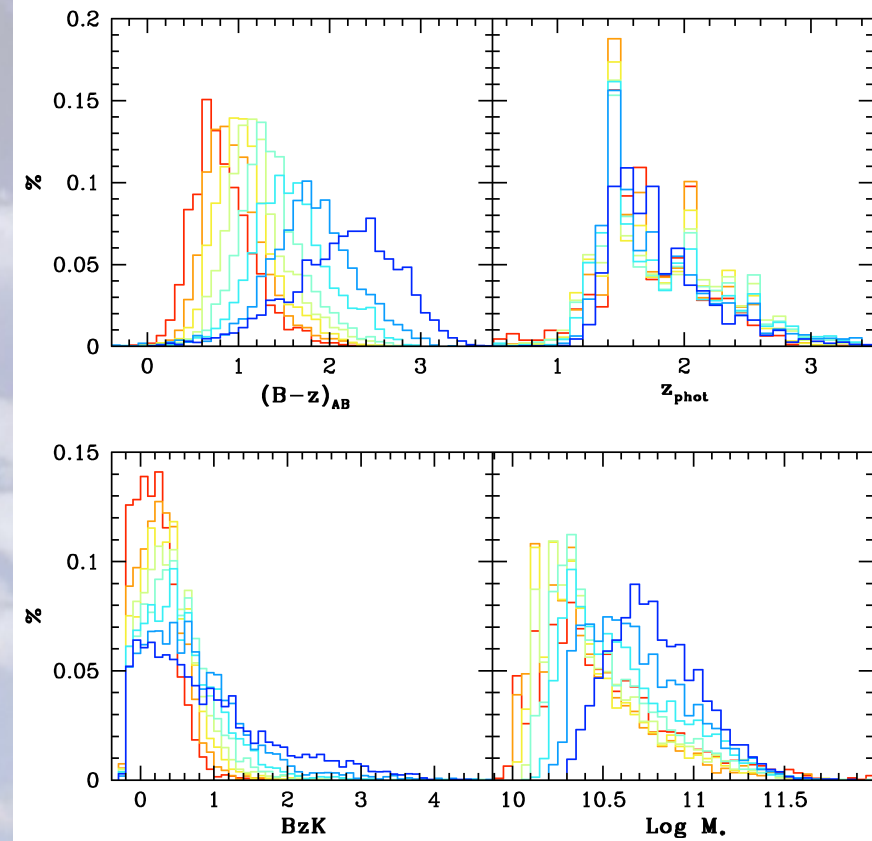
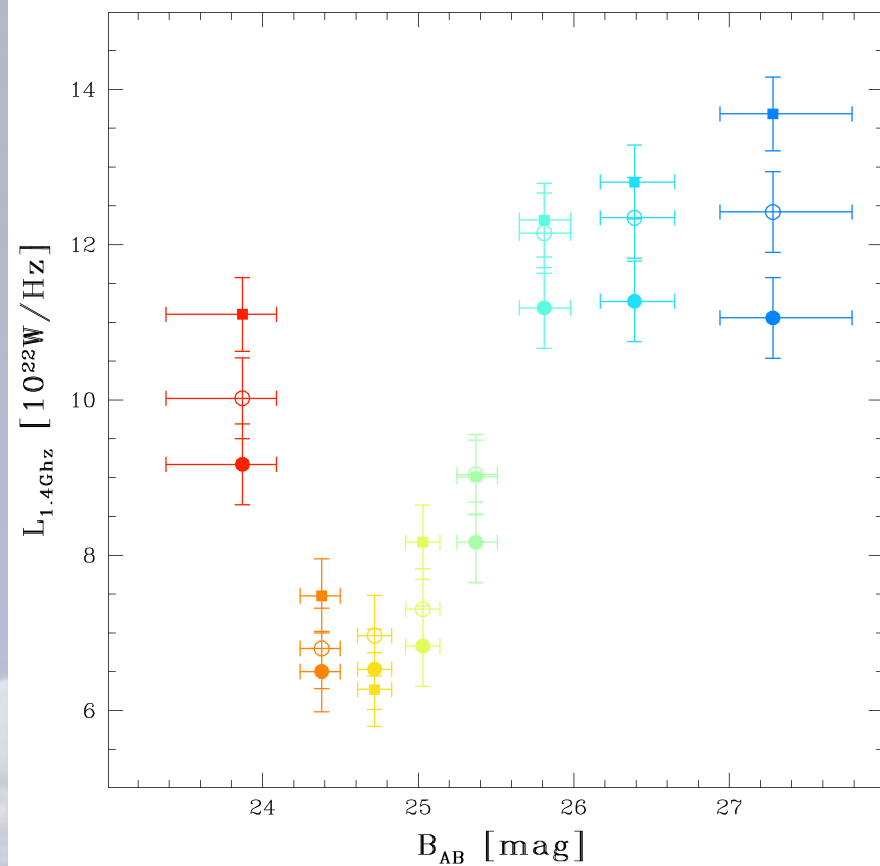
massive galaxies

high star formation rates

# Radio stacks vs. B band mag



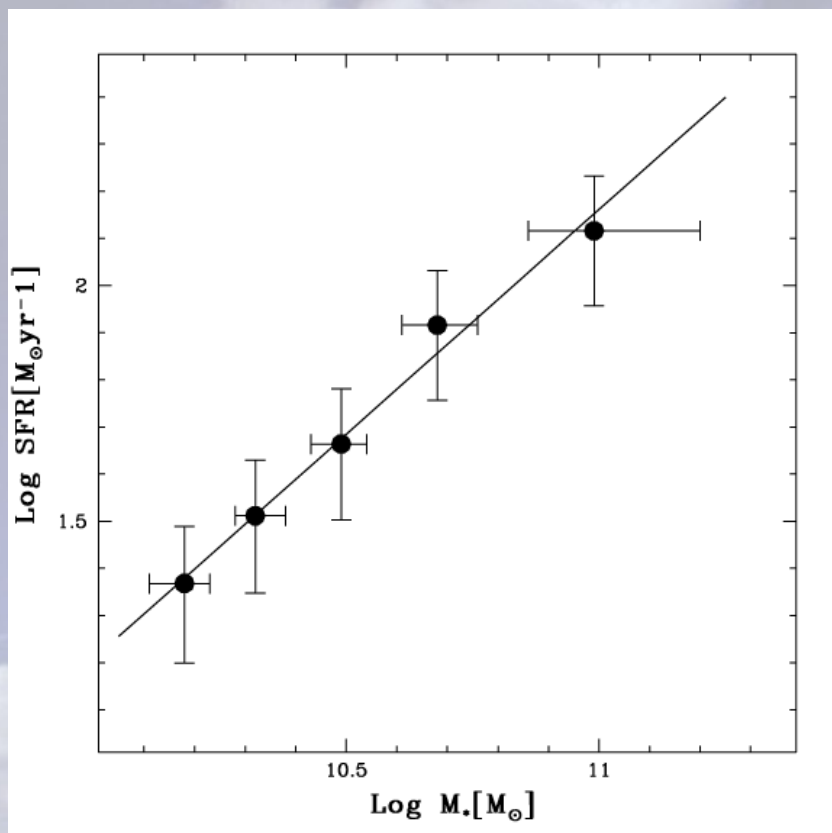
- observed UV restframe light ( $1500 \text{ \AA}$ ) is poorly correlated with the ongoing star formation activity
- counter intuitively: the faintest UV luminosity has the largest SFR





# Radio SFRs vs Stellar Mass

A linear relation is present at all redshifts probed and its slope seems to mildly increase with redshift



**SDSS/AEGIS**

$a = \mathbf{0.7}$  @0/0.7

**GOODS**

$a = \mathbf{0.9}$  @1

Log SFR =  $a \text{Log } M + c$       $a = \mathbf{0.95}$  @1.7

The evolution of the slope sets the time scales of galaxy evolution by tracing when galaxies enter their active stage as a function of mass