

# Noise Performance of ALMA Band10 Receivers Employing High- $j_c$ SIS Mixers

M. Kroug<sup>\*1</sup>, T. Kojima<sup>1</sup>, Y. Fujii<sup>1</sup>, K. Ohtawara<sup>1</sup>, A. Miyachi<sup>1</sup>, and Y. Uzawa<sup>2</sup>

<sup>1</sup> National Astronomical Observatory of Japan, Mitaka, 181-8588, Japan

<sup>2</sup> National Institute of Information and Communications Technology, Koganei, 184-0015, Japan

\*Contact: matthias.kroug@nao.ac.jp

**Abstract**—Current ALMA Band10 cartridges covering the frequency range 790 – 950 GHz are equipped with SIS mixers based on standard Nb/AlO<sub>x</sub>/Nb junctions integrated with NbTiN/SiO<sub>2</sub>/Al strip lines for the embedding circuit. The current density  $j_c$  of the AlO<sub>x</sub>-barrier type junctions is 10 – 14 kA/cm<sup>2</sup>. While the sensitivity performance complies with ALMA receiver noise specifications, tolerance margins are small because of the rather narrow RF bandwidth of the devices [1]. We have fabricated and tested a new set of mixer chips using Nb/AlN<sub>x</sub>/Nb SIS tri-layers with current density  $j_c \sim 30$  kA/cm<sup>2</sup> and low leakage. Circuit design is unchanged except for small modifications in the geometry of the matching circuit to accommodate higher- $j_c$  junctions. Compared with junctions in all-Nb SIS circuits, these junctions usually display a lower gap voltage and higher sub-gap leakage, an observation already made with AlO<sub>x</sub> type junctions. This is likely because of degraded quality of the tri-layer when grown onto the NbTiN film instead of a blank quartz substrate. The high- $j_c$  mixers under test have a gap voltage of  $V_{\text{gap}} = 2.65$  mV and sub-gap to normal state resistance ratios  $q := R_{2\text{mV}}/R_n = 16 - 17$  (all-Nb devices:  $V_{\text{gap}} \sim 2.8$  mV,  $q = 20$  or above). For noise performance characterization mixers are mounted into a Band10 cartridge type test set-up. Measured DSB noise temperatures are below 150 K up to 900 GHz and then increase moderately to  $T_{\text{rx}} = 175$  K at 940 GHz. This fairly flat frequency dependence is a significant improvement over the low- $j_c$  mixers. However, we encountered several difficulties in finding optimum bias conditions for low noise performance *and* linearity. Depending on the position on the mixer's power voltage curve, we observe points with gain expansion or gain compression, a phenomena described in [2]. This is directly related to the voltage swing in the IF output power around  $V_{\text{gap}}/2$  and seems to be more pronounced for mixers based on high- $j_c$  junctions. Details of the measurement method and analysis of the results will be presented at the conference.

[1] Y. Uzawa, "Development and Testing of Band10 Receivers for the ALMA Project," *Physica C*, vol. 494, 2013

[2] C.E. Tong, "Gain Expansion and Compression of SIS Mixers," *IEEE Transactions on Applied Superconductivity*, vol. 19 (3), 2009