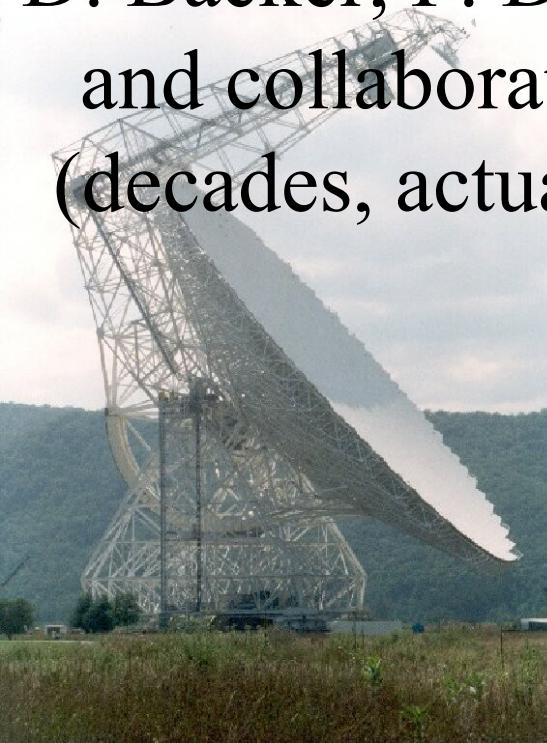


PULSAR TIMING ARRAY: A NANOHERTZ GRAVITATIONAL WAVE TELESCOPE

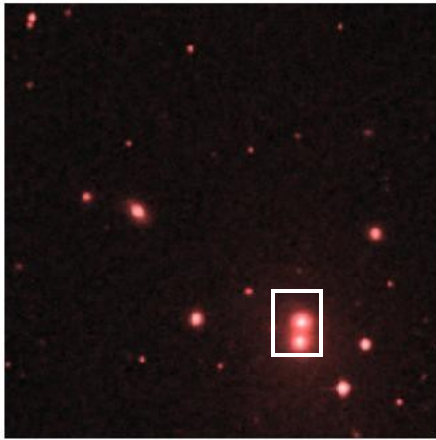
D. Backer, P. Demorest, I. Stairs, R. Ferdman, D. Nice
and collaborators and observatory staffs from years
(decades, actually), NSF funding



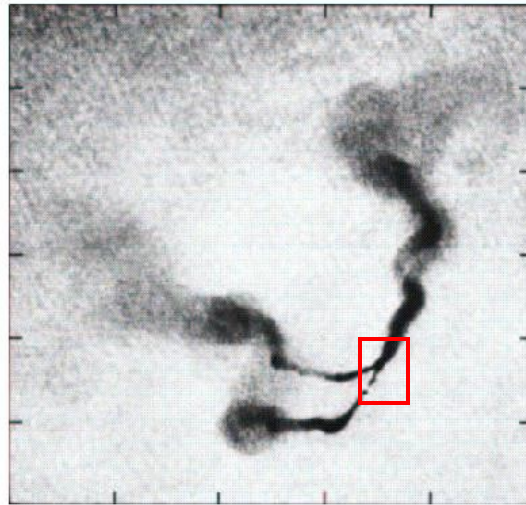
Galaxy Mergers Binary

Massive Black Holes

3C75: Prelude to a Supermassive Black Hole Binary?



Optical image of nuclei
of two galaxies



VLA Radio image of two
jets, and cores

- *Many galaxies contain Massive Black Holes.*
- *All galaxies grow by mergers.*
- When two galaxies with MBHs merge, the two holes will sink to a common center.
- *Compact Massive Black Hole binaries are inevitable!*
- MBH binary orbit will eventually decay via GW emission

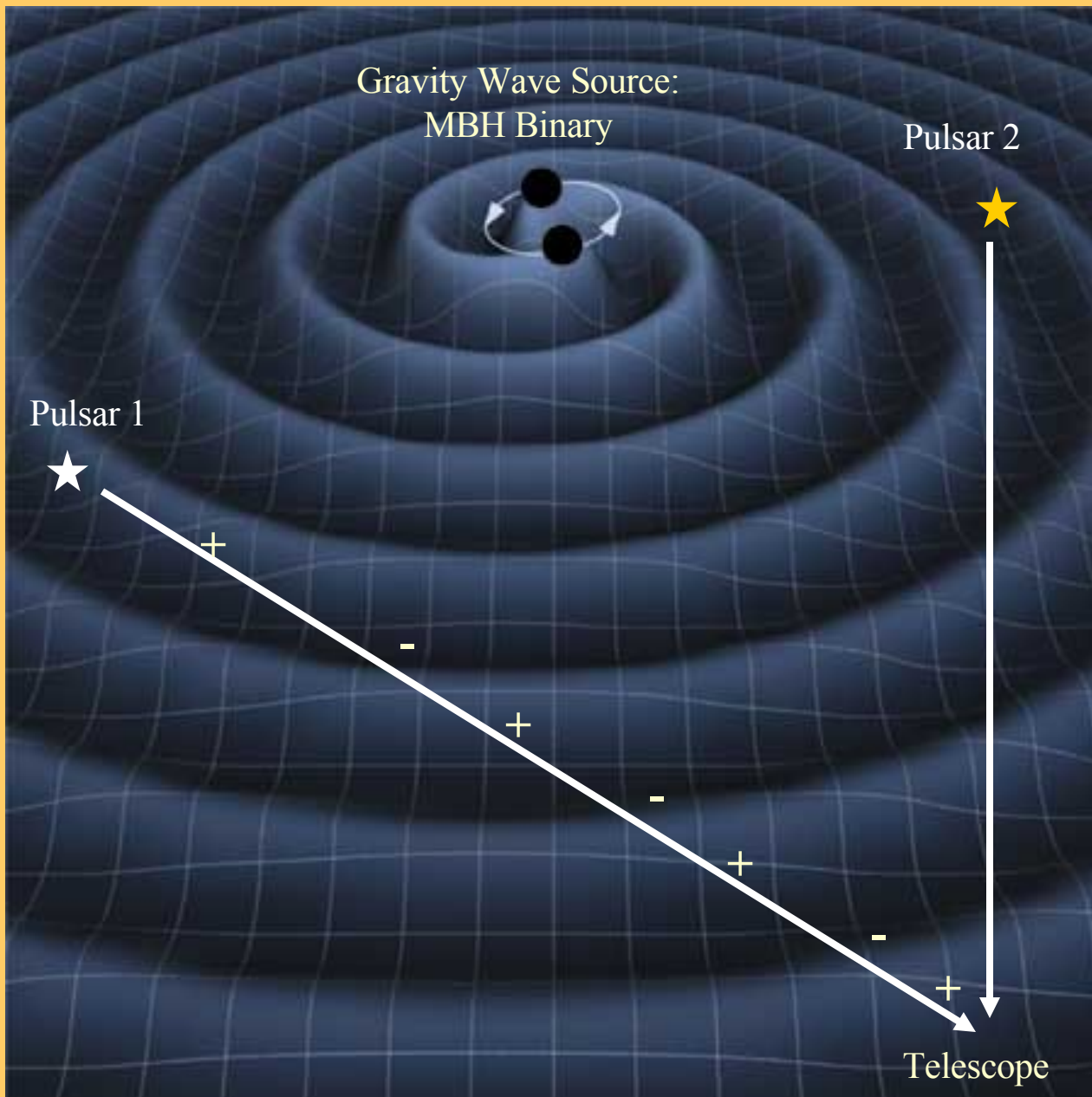
Other Sources

There is renewed interest in cosmic strings resulting from phase transitions in an early Universe with many dimensions (11):

- Damour & Vilenkin (2005) *Phys Rev D*, 71, 063510
- Hogan (2006) *Phys Rev D*, 74, 043526
- Also see Hogan *astro-ph/0608567* “Gravitational Wave Sources from New Physics

Pulsar observations provide the best limit on this fundamental physics phenomenon.

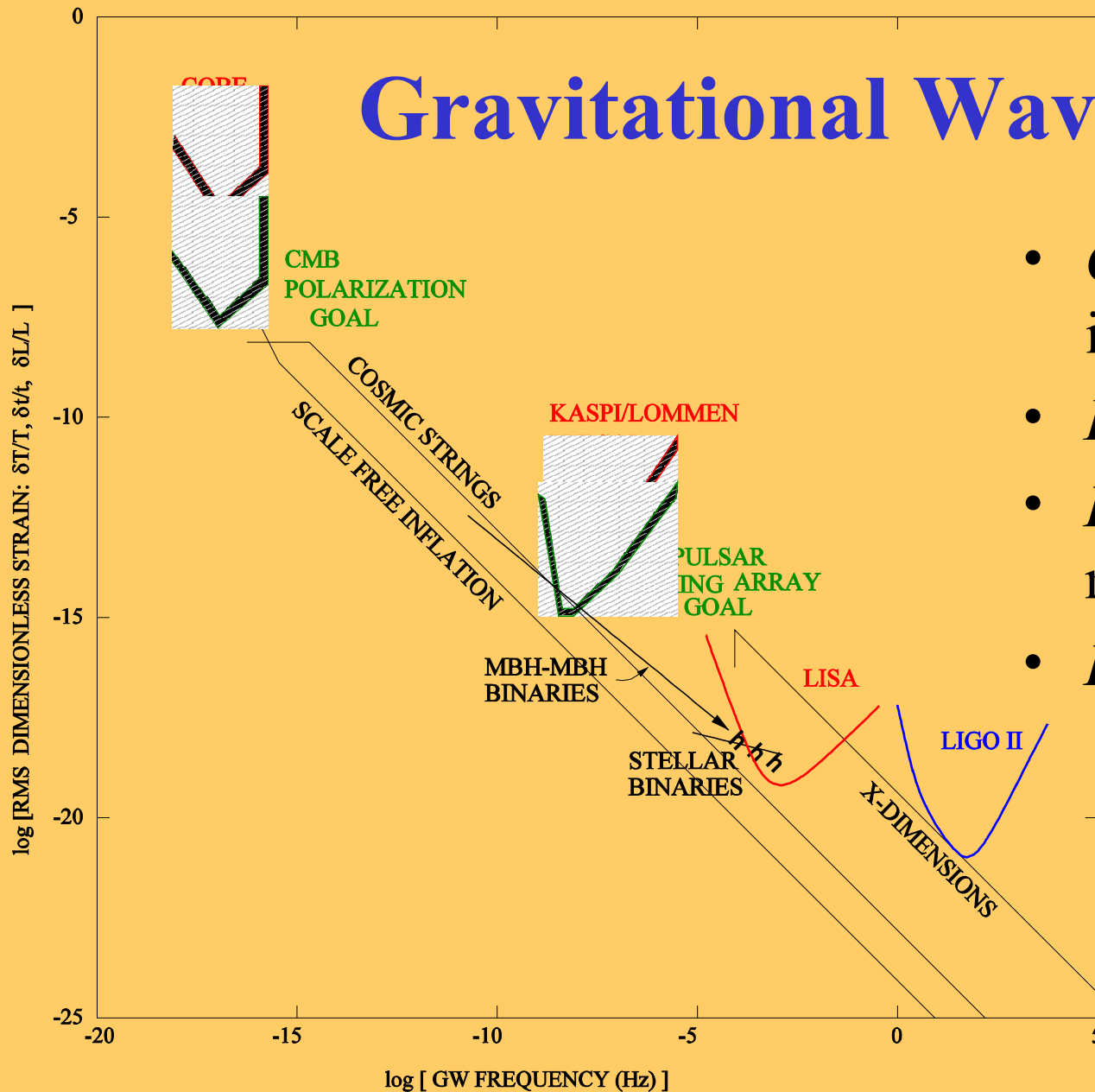
Gravity Wave Detection



$$h_c \sim (rms_{TOA})/T^{1.5}$$

- **Lower rms_{TOA} :**
more “ Bt ” per obs
more/better PSRs
better G/T_{sys}
- **Increase T**

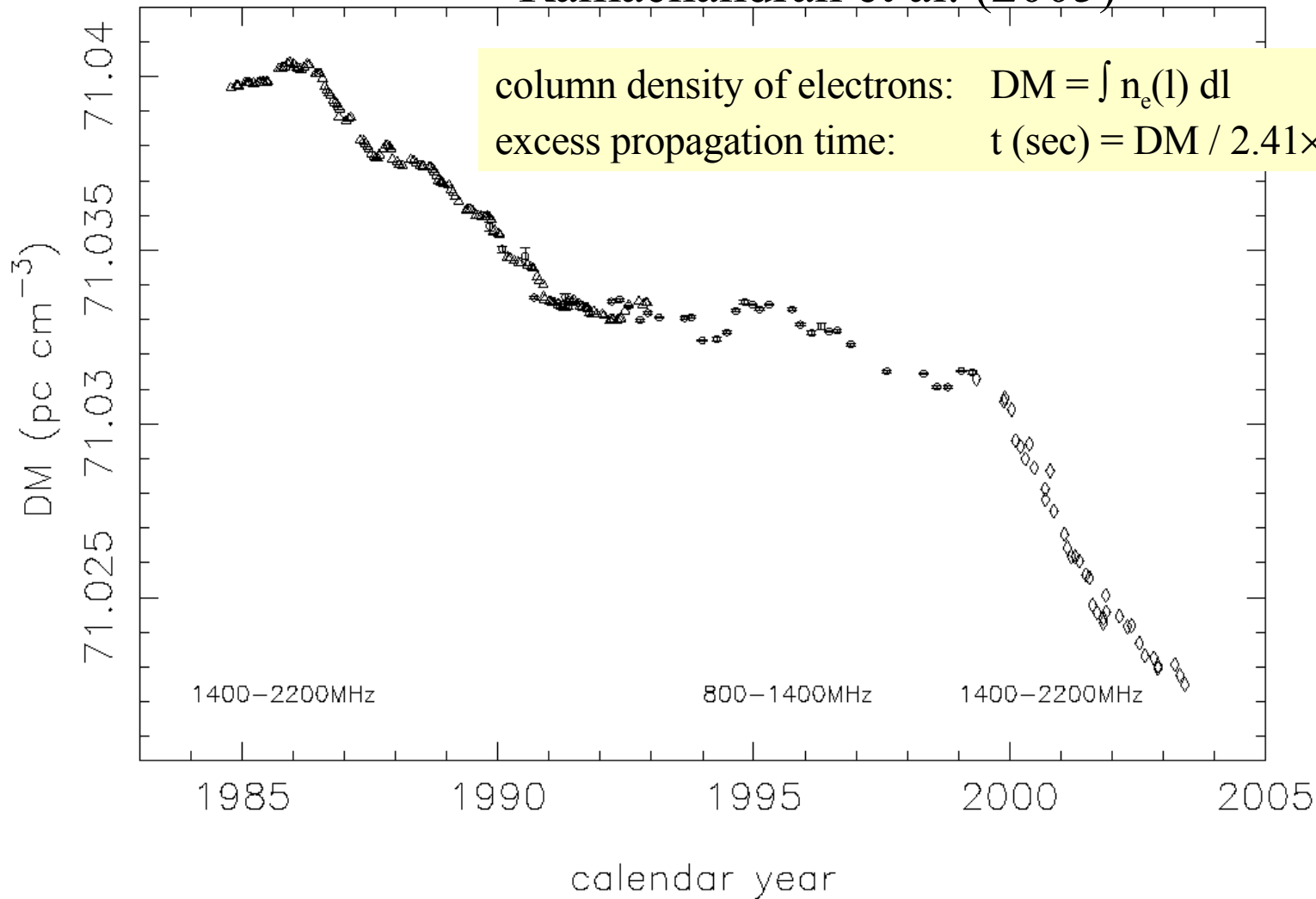
Gravitational Wave Spectrum



- **CMB** Polarization: inflation era probe
- **Pulsars**: MBH-MBH
- **LISA**: MBH-MBH + mBH-MBH
- **LIGO**: Nstar-Nstar

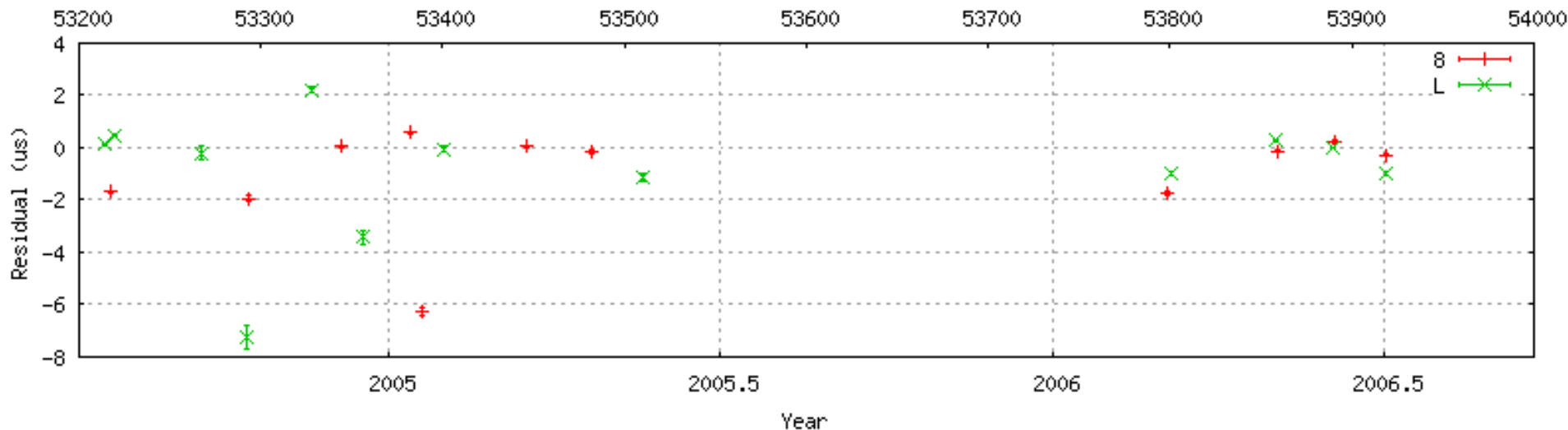
Interstellar Plasma “Weather”

Ramachandran et al. (2005)

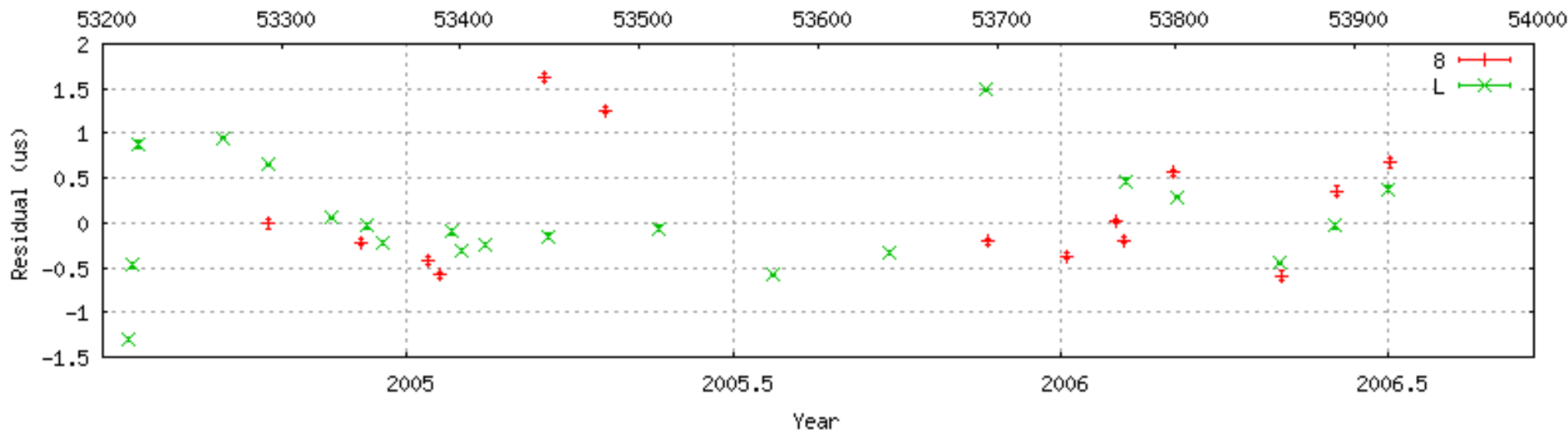


GBT Precision Timing—Demorest PhD

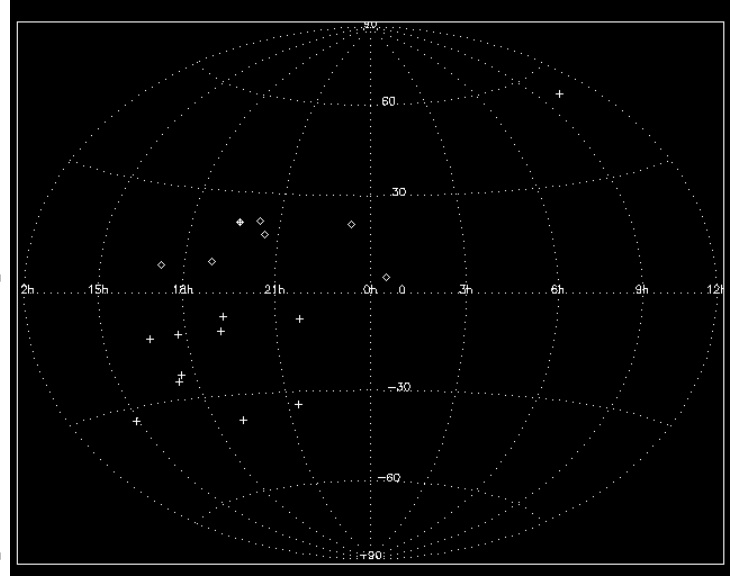
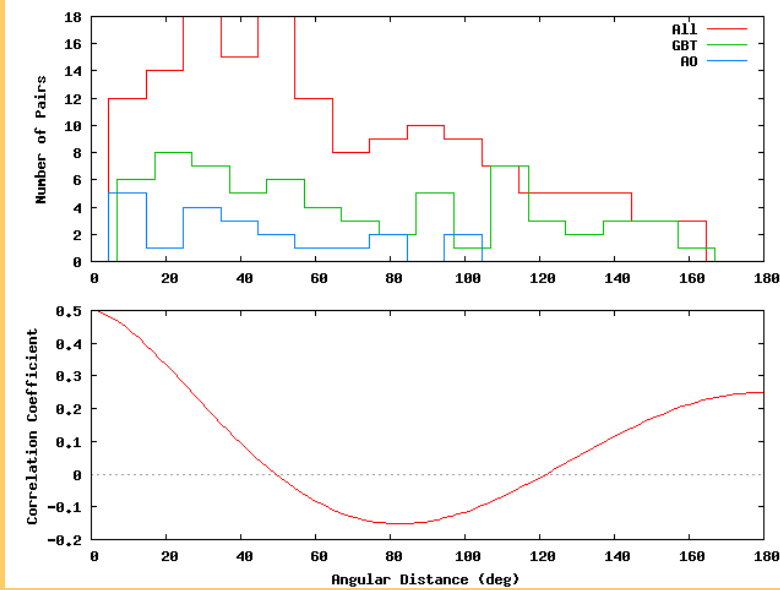
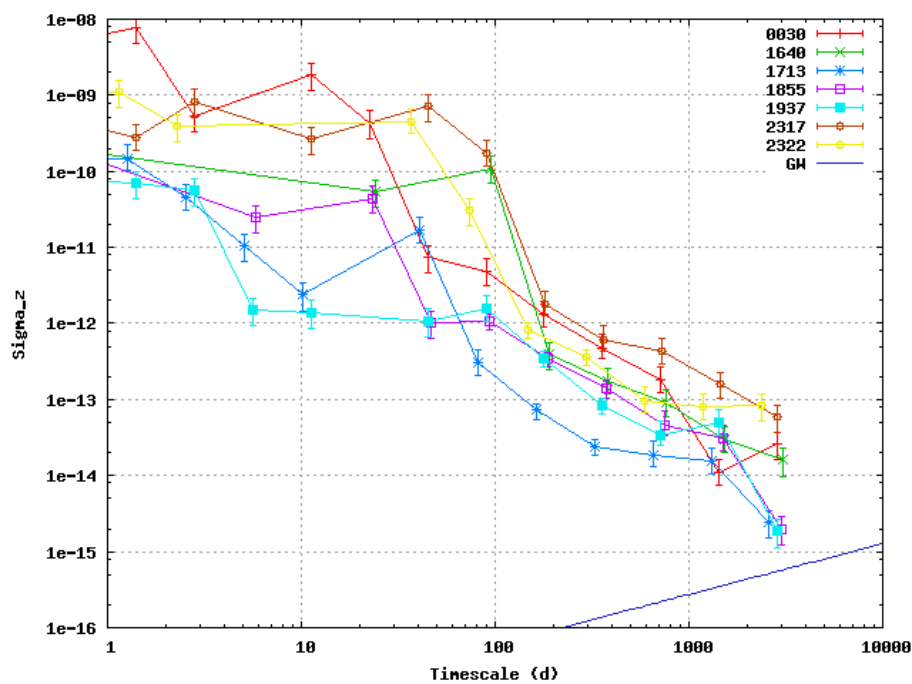
1909–3744 Residuals (1-Day Avg)



1937+21 Residuals (1-Day Avg)

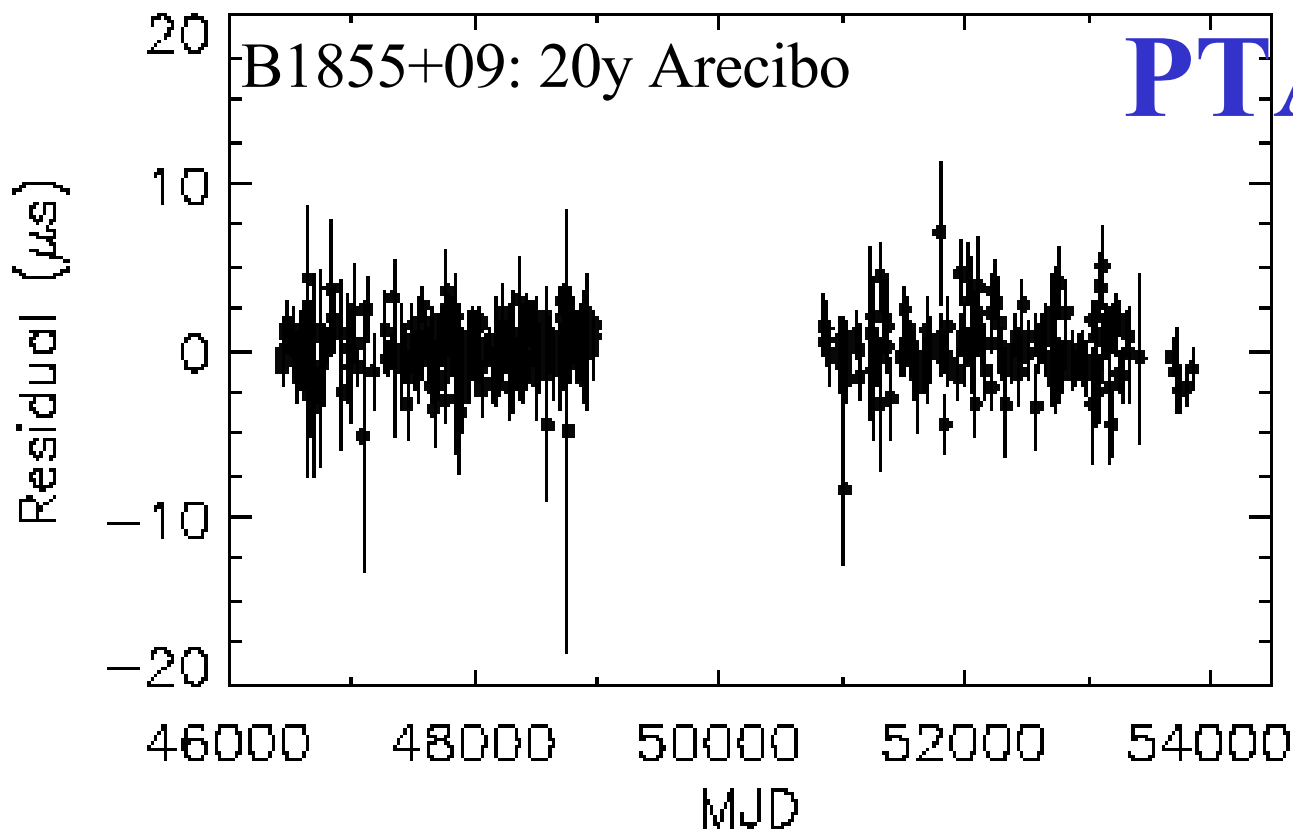


Pulsar Timing Array – Demorest PhD



PTA Results

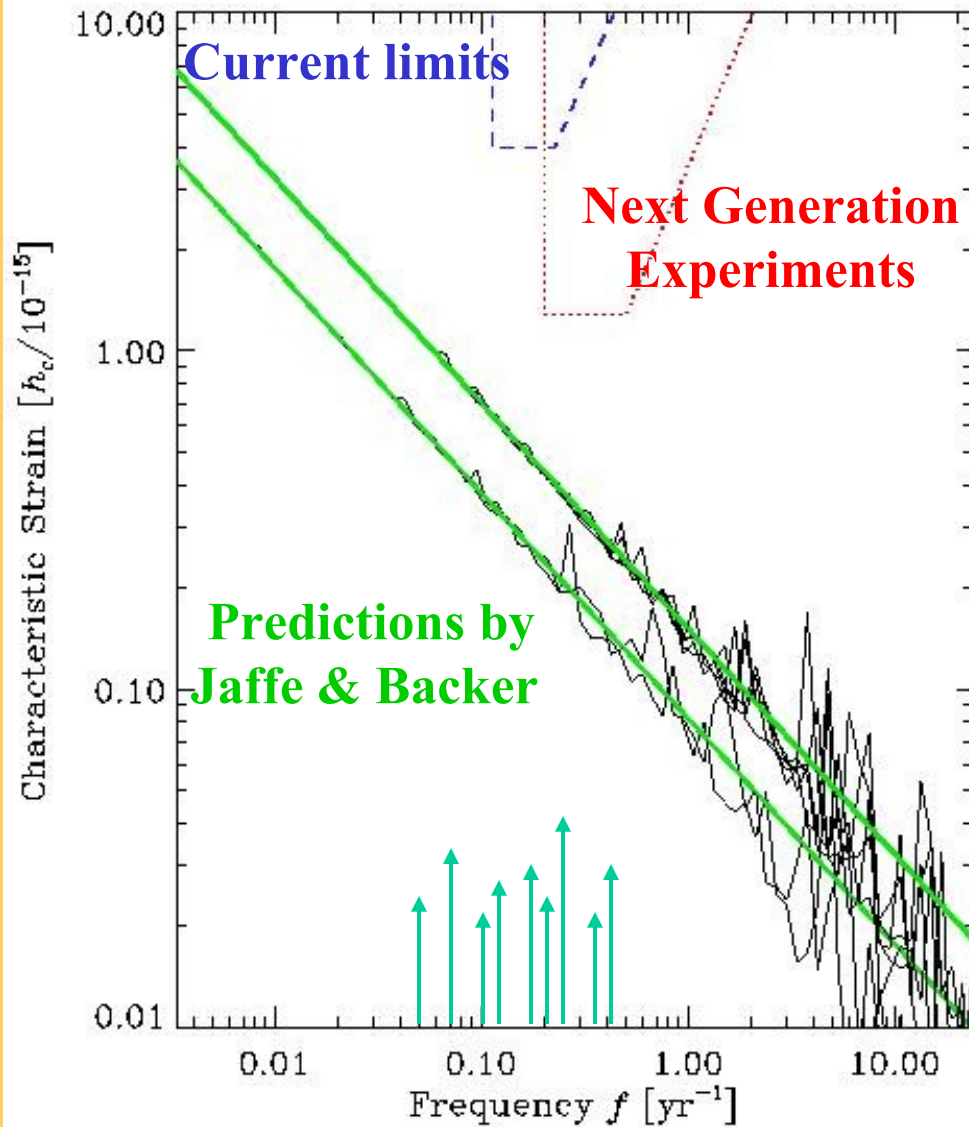
Lommen et al. (2006)



Jenet et al.
(2006)

2006 September 8

Pulsar Timing Limits on the Stochastic Gravitational Wave Background from Massive Black Hole Binaries



Observational Issues

- Sensitivity: $(G/T_{\text{sys}})(Bt)^{0.5}$
- Cadence: monthly (or more often); scheduling important
- Amplitude Calibration: variable gains & polarization apparent profile variations; linearity of system important
- Multi-frequency over more than 2 octaves: DM goes as RF^{-2} ; scattering as RF^{-4} ; $DM(T, RF)$; simultaneous obs important
- Duration: exceed 5y to get beyond 1/y fit “absorption”;
 $h_c \sim (rms_{TOA})/T^{1.5}$
- Phase Calibration: Integrity of clock backend; cable delay from rcvr backend at 10 ns...over 10y
- Size/Integration: large D /small T vs smaller D /larger T ?
- Complementarity of global programs: RF , sky coverage