

Leonid Gurvits, VLBI and Cosmological Parameters

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Theta-z



θ – z relation (Hoyle 1959)



θ – z for extended sources



ng Space VLBI



Angular Size – Redshift Relation for Radio Galaxies

- Tom Legg, 1970 Nature, 226, 64
- Vijay Kapahi, 1987, IAU symposium 124
- Ashok Singal, 1988, MNRAS, 233, 87

Cosmology with compact radio sources

- 1) They are young few hundreds years old short compared to age of universe even at early epoch
- 2) Smaller than host galaxy, they are unaffected by the IGM, or CMB
- 3) Doppler Boosting Oriented close to $1/2\gamma$: Scatter due to projection is small
- 4) Compact sources are mostly quasars; high z



Theta-z



- N= 82 mostly core-jet sources
- λ = 2, 3.8, 6 cm
- $P > 10^{24}$ W/Hz (radio loud)
- DR > 100:1
- Distance to most distant jet > 2%

$$\Lambda = \mathbf{0}$$
$$\mathbf{q}_0 = \frac{1}{2}, \Omega = \mathbf{1}$$

θ -z for compact sources



- N = 337 sources from DSN survey
- 2.3 GHz (13 cm)
- $P > 10^{26} W/Hz$
- Used visibility at longest baseline

$$\Lambda$$
= 0
q₀ = 0.16 ± 0.71
Ω = 0.32 ± 1.42



- N = 350 sources
- 5 GHz (6 cm)
- $L > 2 \times 10^{26} W/Hz$
- -038 < α < 0.18

$$\begin{tabular}{l} \Lambda = 0 \\ $q_0 = 0.21 \pm 0.30$ \\ $\Omega = 0.42 \pm 0.60$ \end{tabular}$$

Angular Velocity – Redshift Relation (μ – z)



Upper envelope consistent with $\Omega = 1$.

$$\beta_{app} = v/c \sim 10$$

Modern VLBI and Cosmology

• VLBI data is much better (VLBA, EVN, S-VLBI)

Ω_m= 0.3

- Cosmology more complex
 - Baryonic Matter
 - Dark Matter
 - Dark Energy $\Omega_{de} = 0.7$
 - $\Omega = \Omega_{\rm b} + \Omega_{\rm dm} + \Omega_{\rm de} = 1$
- Leonid is retired!!





Angular Velocity – Redshift Relation $\Omega = \Omega_{b} + \Omega_{dm} + \Omega_{de} = 1$



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