The Sun Is Hissing at Us

By MARTHA G. MORROW Science Service Astronomy Wester

The sun is hissing at us in a high-pitched voice. And we on earth are just beginning to listen to find out why-and how.

The sun is constantly radiating vast amounts of energy of one sort or another. Light and heat are those with which are most familiar.

As we use higher and higher frequencies in our scramble for radio space, energy reaching the earth in the radio frequency bands becomes of increasing importance.

The general public thru its radio receivers makes use of radio waves about a quarter of a mile long, and cover: radio hams communicate on waves about the length of a city lot; experimental scientists are using still shorter waves to explore the sun.

STEADY HISS

Solar noise, which interferes with radio reception at ultra-high fre-quencies, is a steady hiss. Upon this are often superimposed "puffs" and "swishes" lasting but a second or less. When the swishes overlap, a grinding noise results.

a grinding noise results.

Solar static comes to us on the sama frequencies as those used for FM, television and radar. Giant "mirrors" 10 to 25 feet across are being installed to explore outbursts from the sun.

Two of these enormous mushrooms are being set up at the National Bureau of Standards' radio propagation laboratory at Sterling, Va. Originally Ciant Wurzburgs, a type of radar used by the Germans, these instruments have been converted to record solar noise. Brought back to this country by the Army Signal Corps, their steel mesh mirrors are 25 feet across.

SECOND NEARLY READY

SECOND NEARLY READY

A small antenna in the center of the basket-shaped mirror receives solar static, which is carried by cable to the electronic equipment in an adjoining building. One saucer is already at work on the 480 to 500 megacycle band; the other will begin collecting static within a few months.

Standards' project, hopes to dis-

- · Frequencies on which the solar. noise is strongest.
- If the static varies with the seasons.
- . Whether or not there is a longterm fluctuation.

MORE THAN WE SEE

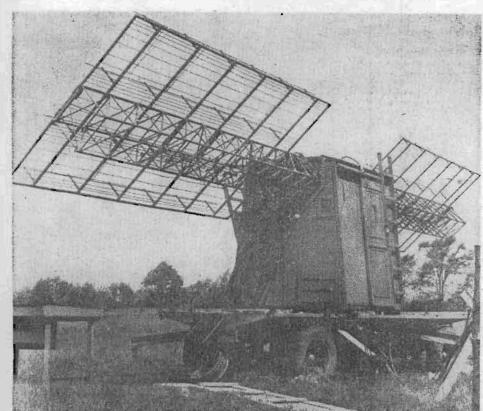
edge to be seen visually.

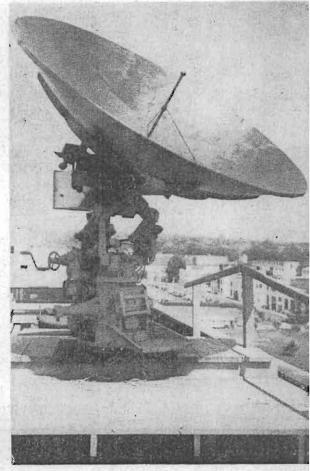
Thus we are beginning to learn much about the sun, source of our heat, light and other energy. Hisses from the sun are attracting the attention of an ever-increasing numan adjoining building. One salter is already at work on the 480 to 500 megacycle band; the other will begin collecting static within a few months.

With these instruments Grote Reber, in charge of the Bureau of surface or corona. Thus science has means of exploring our nearest star.

Glant radar mirror, above, at the Bureau of Standards radio propag tion laboratory at Sterling, Va., is used for study of ultra-high fr quency radio noise generated by the sun.







Solid metal dises, above, are used at the Naval Research Laborator Anacostia, to collect the sun's radio waves.

A "mirror" is needed for the mount, at left, before Cornell Uc versity's "radio telescope" will be complete.