Subject: [allemploy] FYI: 21 Nov BIWEEKLY CALENDAR OF THE ALMA PROJECT at NRAO From: Al Wootten <awootten@nrao.edu> Date: 11/22/2005, 5:14 PM **To:** alma-info@nrao.edu, anasac@nrao.edu, allemploy@nrao.edu BIWEEKLY CALENDAR OF THE ALMA PROJECT at NRAO Nov 21, 2005 -- Dec 5, 2005 Happy Thanksgiving to All! A scientific Meeting was held on 'Physique stellaire avec ALMA' November 14/15, 2005, GRAAL - Montpellier, France. Jerome Pety reports considerable interest in ALMA in this community, with interesting discussions of science goals achievable with the array. Past issues of this Calendar may be viewed at http://www.cv.nrao.edu/~awootten/mmaimcal/ALMACalendars.html See also the JAO ALMA Calendar overview at: http://www.alma.cl/alma project General Happenings OSF: An anniversary celebration will be held in San Pedro de Atacama to celebrate the community's founding. Tender documents for grading of the Vertex contractor area have been released. Bidding process for OSF, AOS Technical Building finishes, camp services, and road finish (culverts, arroyo crossings etc) are all at various levels of progress. Rough work on road in progress at km 40. There are about 125 persons working at the site. AOS: Construction continues on AOS Technical Building shell. ASTE reports no damage from the northern Chile earthquake last week. ATF: Preparations continue for optical pointing tests. The AEC antenna has moved. AOC: Two line length corrector (LLC) modules and one laser synthesizer module delivered to prototype system integration (PSI). Antenna to AOS Technical Building simulation test completed with excellent results. NTC: Beam patterns measured on B6 (1.3mm) cartridge no. 2 show significant improvement. Construction of cartridges 4 and 5 under way. Correlator Quadrant No. 2 bin and power rack assemblies completed. First tunable filter bank (TFB) card with final components passed tests and was shipped. NAASC: Computing IPT Leads meeting was held in Charlottesville, along with meeting of the Science Software Requirements committee. The program for the Z-Machines meeting has been published on the meeting website http://www.cv.nrao.edu/naasc/zmachines. DAILY CALENDAR (Times EDT ) see https://wikio.nrao.edu/bin/view/ALMA/AlmaCalendar Mon 21 November All Day: SSR Meeting, Charlottesville Tue 22

10:30 AM-11:30 AM: JAO IPT Telecon 4:00 PM-5:00 PM: NAScienceIPT teleconference (open to all interested parties) (434)296-7082 Wed 23 Thu 24 Thanksgiving Holiday, U. S. Fri 25 NRAO Holiday, U. S. Sat Nov 26 Sun Nov 27 Mon Nov 28 Tue Nov 29 10:30 AM-11:30 AM: JAO IPT Telecon 4:00 PM-5:00 PM: NAScienceIPT teleconference (open to all interested parties) (434)296-7082 Wed Nov 30 Thu Dec 1 Fri Dec 2 Sat Dec 3 Sun Dec 4 \* Nov 21-22 -- SSR Meeting, Charlottesville \* Nov 30-Dec 1 -- National Science Board meets \* Dec 7-8 -- ESO Council Regular Meeting \* Dec 7 -- ASAC Telecon \* Dec 7 -- ANASAC Telecon \* Dec 22 -- ALMA Board Telecon \* Jan 4-7 -- URSI/NA, Boulder \* Jan 12-14 -- ALMA Zmachines workshop ALMA Memo 543: Estimating Calibrator Counts at 250 GHz Using MAMBO Observations of Flat Spectrum Quasars Author: M.A. Holdaway, Chris Carilli, Axel Weiss, Frank Bertoldi Abstract: We analyze the 250 GHz MAMBO fluxes of cm-selected bright, compact, flat spectrum quasars, match them with 8.4 GHz CLASS flux measurements, and derive a distribution for the spectral index between 8 GHz and 250 GHz. This spectral index distribution, when combined with Condon's 5 GHz flat spectrum source counts and the distribution of core fraction taken from the flat spectrum members of the 3CR2 sample, provides us with an estimate of the source counts of bright, compact, flat spectrum quasars which will be available to ALMA at 250 GHz for various calibrations. Over the entire sky at 250 GHz, we find there should be about 28,000 quasars brighter than 10 mJy, 2230 quasars brighter than 100 mJy, and 70 quasars brighter than 1 Jy. The source count estimates in the current memo exceed the estimates of Holdaway and Owen (2005) at 250 GHz by 14% at 10 mJy, 33% at 100 mJy, and 55% at 1 Jy. The higher estimated counts as derived in this current work are a direct result of sources observed to have a very flat or even inverted spectral index between 8 and 250 GHz which were largely missing in Holdaway and Owen's spectral index distribution. A pdf copy of this memo is on edm.alma.cl and at: http://www.alma.nrao.edu/memos/html-memos/alma543/memo543.pdf \_\_\_\_\_ ALMA Memo 544: Quasi-Optical Verification of the Band 9 ALMA Front-End Author: M. Candotti, A. M. Baryshev, N. A. Trappe, R. Hesper, J. A. Murphy, J. Barkhof, W. Wild

Abstract: The front-end optical design for band 9 (600 to 720GHz) of the Atacama Large Millimeter Array (ALMA) is now completed and verified. A frequency

independent design approach is used to couple radiation to the two orthogonal polarized mixed detectors from the large 12m ALMA Cassegrain telescope. As it is a heterodyne receiver, two local oscillator beam paths are integrated into the front-end optical system. Due to the large number of interferometer elements (64 antenna units) to be built, installed and maintained in the remote site of the Atacama Desert, reliability of the optical system should be ensured. A modular and compact optical design is also important. In addition a cheaper fabrication process is considered, at these more tolerant higher frequencies, by milling the mirror surfaces near the surface roughness limit. In this paper we verify the optical design and estimate system efficiency by means of experimental measurement and software simulation comparisons. Precision planar scans of near field beam patterns (amplitude and phase) have been measured. Experimental beam measurements were taken at the output of the mirror coupling system (telescope focal plane location) for both polarization paths and for both local oscillator beam guides. At the same measurement locations, software simulations of a highly accurate geometrical model of the mirror coupling system were predicted using the commercial package GRASP8. These comparisons at some fundamental locations along the beam paths, allow the assessment of the quasi-optical beam coupling system design. The local oscillator power budget analysis is carried out from results obtained using GRASP8. In the conclusion we summarize the current status and describe future analysis plans.

Construction

The National Radio Astronomy Observatory expects to invite applications for a NRAO Postdoctoral appointment at the North American ALMA Science Center in Charlottesville, Va. These positions provide 50% time for independent research, with the remaining 50% assigned to project and operational duties at the sites.

The position is funded by a grant from the National Science Foundation through the ALMA Construction Project. See

Please send information for upcoming calendars by Friday evening of the preceding biweekly period to Jennifer Neighbours or Al Wootten via e-mail (jneighbo@nrao.edu or awootten@nrao.edu).

The calendar will be issued between late Friday and sometime on Monday by e-mail to all NRAO scientific staff members and anyone else interested. A specific mailing list, alma-info, has been created for anyone wishing to receive it. Past issues are available at http://www.cv.nrao.edu/~awootten/mmaimcal/ALMACalendars.html [allemploy] FYI: 21 Nov BIWEEKLY CALENDAR OF THE ALMA PROJE...

Allemploy mailing list Allemploy@listmgr.cv.nrao.edu http://listmgr.cv.nrao.edu/mailman/listinfo/allemploy