## Developments in Heterogeneous Computing at Green Bank

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## Heterogeneous Computing Systems

A system made up of architecturally diverse userprogrammable computing units.

Typically a mixture of two or more of the following:

Traditional RISC/CISC CPUs

Graphical Processing Units (GPU)

Reconfigurable Computing Elements (RCE)

## **Real-time Computing**

A real-time computing system must operate while meeting fixed deadlines for completion of scheduled computations

Computations finished correctly but late are not useful

Real-time does not mean "really fast"! But it is helpful!

## Heterogeneous Real-time Computing

Combines both ideas into one system

Note that many heterogeneous systems are not designed to be real-time systems:

SRC-{6,7}, Convey, CPU/GPU clusters, supercomputers

Building a heterogeneous system is more difficult than building a homogenous system due to the need to master multiple toolsets, computing models, and hardware characteristics.

Add in the requirement for real-time response, and it is even more challenging.

# Then Why Bother?

- The data rates are too high for a general-purpose machine
  - Use an FPGA preprocessor to operate on the raw data streams for I/O management
  - Use a GPU to offload computations from the CPU for CPU load management
- A general purpose machine would demand too much power
  - FPGA's and GPU's are more efficient at computations that fit their abilities

## **Computer Systems Architecture**

**Processor architectures** 

CPU

X86\_64 – Intel and AMD

Everything else – IBM Power and Cell

GPU

NVidia

AMD/ATI

Reconfigurable Computing Elements (FPGA-based processors)

## **Reconfigurable Computing Elements**



### Pulsars



## An Example: GUPPI

Pulsar "Search" backends:

Fast-dump spectrometers

Time resolution 50 to 100 µs

Frequency resolution 25 kHz to 1 MHz

Often trade data quality for more BW

Pulsar "Timing" backends:

Pulse period folding

Coherent dedispersion

High time resolution,  $\sim 1 \ \mu s$ 

Data quality (# bits, polns, etc) more important

## **GUPPI Block Diagram**









#### GUPPI architecture: ~1 MHz PFB in FPGAs Coherent dedisp in GPUs





10 Ge 24 Gb/s



P. Demorest

>10x improvement in BW!

Fully utilizes all GBT lowfreq receivers.

Improved S/N ratio, also reduces scintillation variability.

~ 350 MB/sec sustained data rate (~10 TB/8 hours)

PSR J1713+0747 plot: S. Ransom



### **Profile comparisons**



### Conclusions

### Heterogeneous computing works.

- FPGAs for the fastest streaming integer signal processing
- GPUs for fast floating point DSP calculations
- General-purpose CPUs for managing the system, assembling and transmitting data, and for parts of the computations that are not amenable to the GPU

## New GBT Spectrometer

- Joint effort with UC Berkeley NSF ATI project
- 16 IF inputs
  - 8 dual-pol beams, or 16 single-pol beams
- 3 GS/s sampling rate @ 8 bits/sample
- Can be ganged to achieve ~10 GHz instantaneous bandwidth on 2 polarizations



### **Spectrometer Data Rates**

- ADC to FPGA: 3GS/s \* 16 Polns \* 8 bits = 384 Gb/sec
- FPGA to GPU: ~9 Gb/sec \* 8 GPUs = 72 Gb/s

- Raw (mostly) Data Output Rate:
- Maximum specified at 33 MB/s \* 16 polns = 525 MB/s, or 15 TB/8 hours, or 1 PB every 66 hours.
- Front-end (FPGA-GPU) hardware is capable of ~10x this rate...

## Conclusions

Even today:

We can collect data faster than we can write it to disk

We can collect data faster than humans can monitor it for quality

We can't afford to store raw data even @ the raw disk drive price of \$90K/PB.

Our RF bandwidth is >4x our digitized bandwidth



# Summary of Required Observing

TIL A CDT

Table 2: GBT spectrometer modes specified per beam (2 IFs) <sup>w</sup>							
Number of	Sub-band	Number of	Spectral	Velocity	Velocity	Integration	
sub-bands	$Bandwidth^b$	channels per	resolution	range at	resolution	$\operatorname{time}$	
per IF		sub-band per IF		$90  \mathrm{GHz}$	at 90 GHz	minimum	maximum
	(MHz)	_	(KHz)	$(\mathrm{km~s^{-1}})$	$(\mathrm{km}\;\mathrm{s}^{-1}$ )	(msec)	(sec)
Observing Mode 1							
1	$1500^{c}$	1024	1465	5000 <sup>b</sup>	4.9	0.5	60
1	1000	2048	488	3333	1.6	0.7	60
1	800	4096	195	2667	0.7	1.3	60
1	500	8192	61	1668	0.2	2.5	60
1	400	16384	24	1333	0.08	5	60
1	250	32768	7.6	833	0.03	10	60
1	100	32768	3.1	333	0.01	10	60
1	50	32768	1.5	166	0.005	10	60
1	25	32768	0.8	83	0.003	10	60
1	10	32768	0.3	33	0.001	10	60
1	5	32768	0.15	17	0.0005	10	60
1	1	32768	0.03	3	0.0001	10	60
Observing Mode 2							
8	30	4096	7.3	100	0.02	10	60
8	15	4096	3.7	50	0.01	10	60
8	10	4096	2.4	33	0.008	10	60
8	5	4096	1.2	17	0.004	10	60
8	1	4096	0.2	3	0.0008	10	60

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 $^{a}$  These modes are implemented in each spectrometer that processes 2 IFs from a beam.

<sup>b</sup> In Observing Mode 1, bandwidths less than 1500 MHz should be centered between 150 MHz and 1350 MHz.

 $^{c}$  The usable bandwidth will be 1250 MHz, which corresponds to a velocity range of 4165 km s<sup>-1</sup> at 90 GHz.



#### Exa-problems for SKA computing

- Timescale 2014 +
  - 100% science observations ~ 2020
- Management of complexity
  - · Many stake-holders and users
- Networking
  - ~ 1Tb/s per antennas short haul (< 200km)</li>
  - 160Gb/s per station long haul (> 200km)
- Data flow
  - · Data rate ~ 1 EB/day
  - Real time processing vital!
- Parallel processing
  - 1 EFlop/s ~ 1 billion cores
  - · How do we program these?
- Computing power requirements
  - · 1Eflop/s = 350MW at Blue Gene efficiency
  - · Need one to two orders of magnitude improvement
  - Plus green power
- Minimize use of radio-astronomy specific technology

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