

# Terahertz MgB<sub>2</sub> HEB mixers with a 13GHz gain bandwidth

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Broad intermediate frequency (IF) bandwidth is a crucial feature for heterodyne receivers utilized in terahertz astronomy [1]. Whereas below 1.2 THz superconductor-insulator- superconductor (SIS) mixers have demonstrated an IF bandwidth in excess of 20GHz , at frequencies > 1.2THz superconducting hot-electron bolometer (HEB) mixers have not come to such level of performance yet.

We believe that utilizing MgB<sub>2</sub> ultra-thin [2] films the IF bandwidth for HEB mixers can be extended way above 20 GHz, hence satisfying the most demanding astronomical tasks. The key for a large IF bandwidth in HEB mixers is fast electron- phonon interaction and short phonon escape time from the superconducting film into the substrate. Recently we have shown that MgB<sub>2</sub> films as thin as 5nm can be made on SiC substrates, with a critical temperature of 30K. The minimum receiver noise temperature has been shown to be at 1000K (1.6THz) [3]. However, though the demonstrated IF bandwidth was 11GHz, we had some reasons to believe that this is far to be the limit.

By performing mixing experiments in a broadband cryogenic probe station we have been able to show that the gain bandwidth (3dB gain roll-off) in our MgB<sub>2</sub> HEB mixers is 13-14GHz, whereas the mixer noise temperature was

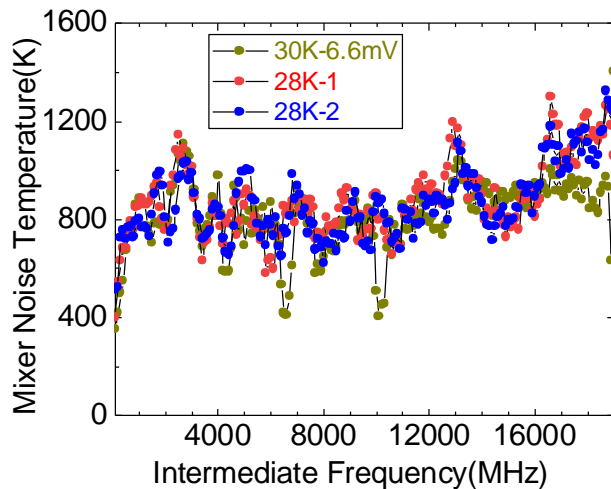


Fig. 1. The MgB<sub>2</sub> HEB mixer noise temperature spectrum.

nearly constant up to an intermediate frequency of 20GHz (Fig.1). These experiments were conducted at a rather low LO frequency of 100GHz, yet results have clearly shown that the full potential of MgB<sub>2</sub> HEB mixers is far not reached.

In order to justify applicability of low LO frequency experiments for modeling the high LO frequency  $\nu(h\nu > 2\Delta$ , where  $\Delta$  is the superconducting energy gap) we designed a set of experiments where MgB<sub>2</sub> HEB mixer performance (gain, noise, IF bandwidth) is studied vs LO frequency and operation temperature (which affects the superconducting energy gap  $\Delta$ ).  $\Delta$  can be obtained from the measured kinetic inductance (from 5K and up to the critical temperature  $T_c$ ).

## REFERENCES

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